



# HXMT views on the accreting millisecond X-ray pulsars

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FPS12, 2023. 7. 6

# Outline

- Recycling scenario
- AXMPs
- Swift J1756.9-2508 & MAXI J1816-195
- Summary & Future

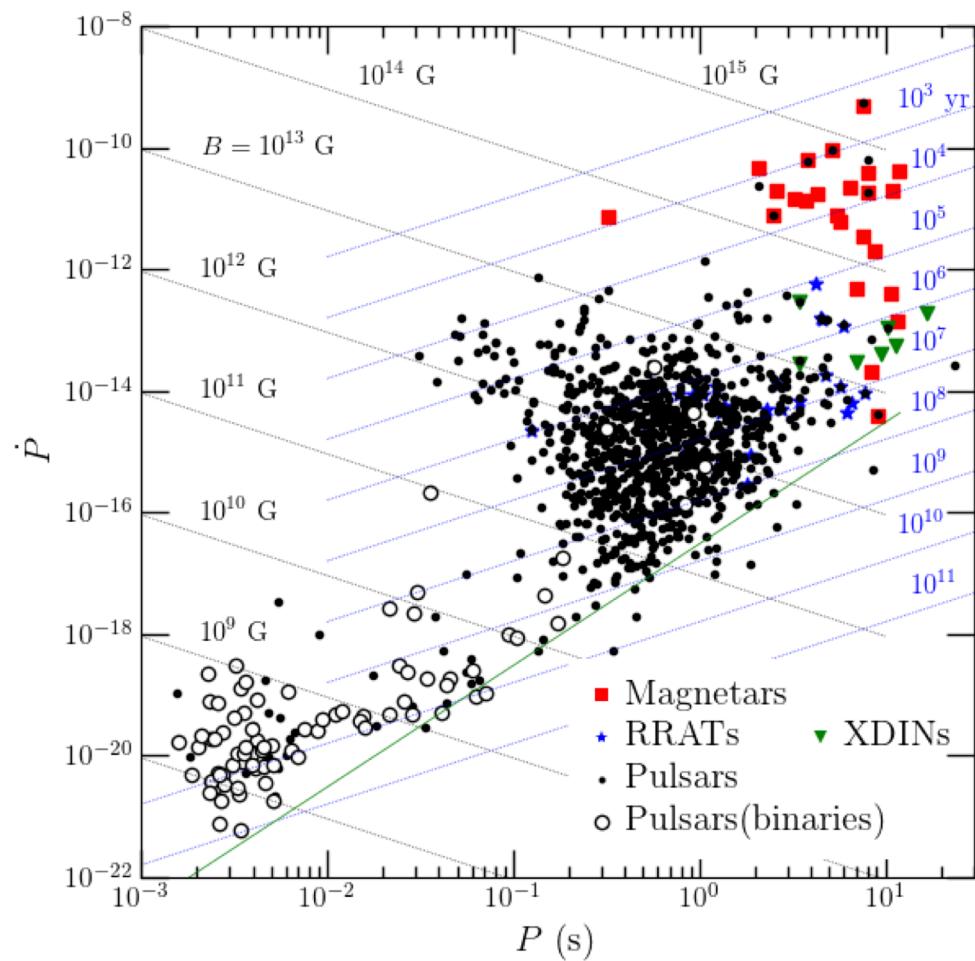
# Recycling scenario

- 1982, the first radio millisecond pulsar, PSR B1937+21

Alpar et al. 1982

Radhakrishnan

& Srinivasan 1982



# Recycling scenario

**1996, burst oscillation, 4U 1728–34, ~363 Hz**

**1998, SAX J1808. 4–3658, pulsation, ~401 Hz**

**2003, SAX J1808. 4–3658, burst oscillation, ~401 Hz**

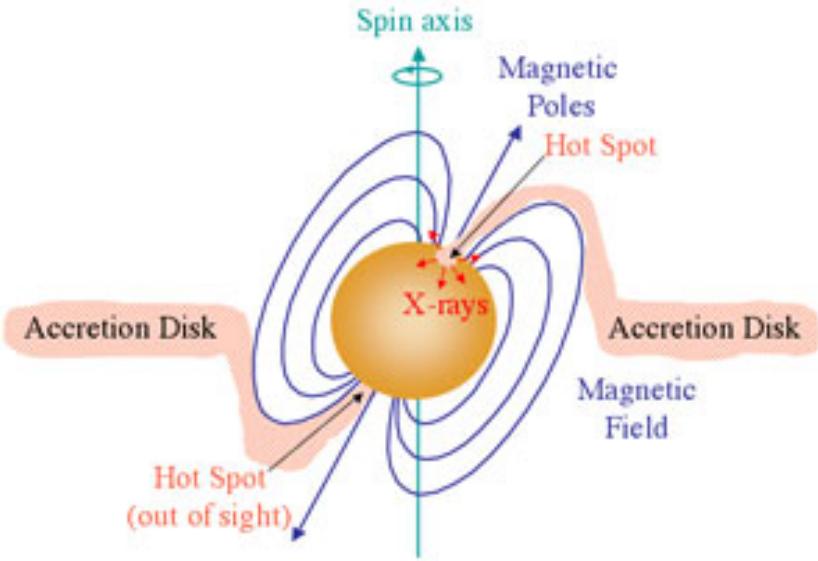
**2005, spin-up, IGR J00291+5934**

**2013, switch between accretion and rotation power**

**IGR J18245–2452, PSR J1023+0038, XSS J12270–4859**

# Accreting millisecond X-ray pulsar

- Subgroup of LMXBs
- 23 AXMPs confirmed
- $P_s \sim 1.7 - 9.5$  ms
- $P_{orb} \sim 40$  min – 11 hr
- Four AMXPs observed by HXMT

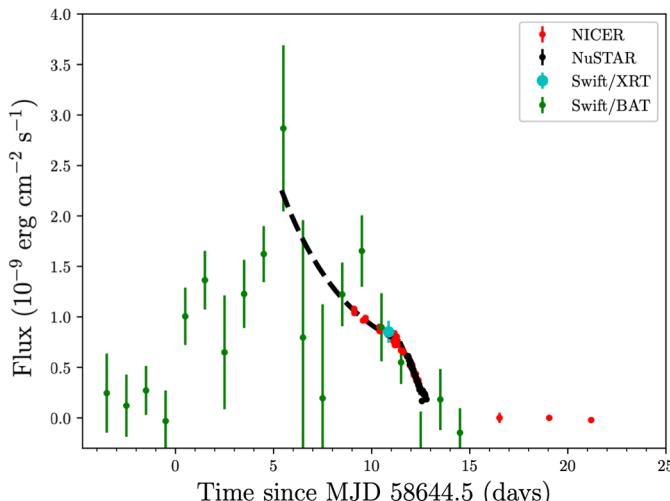
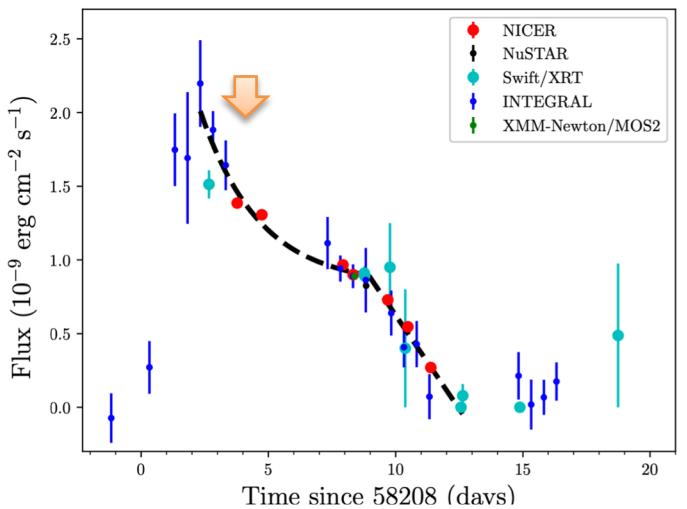


Swift J1756.9-2508 & MAXI J1816-195

SAX J1808.4-3658 & IGR J17498-2921

# X-ray pulsation: Swift J1756.9-2508

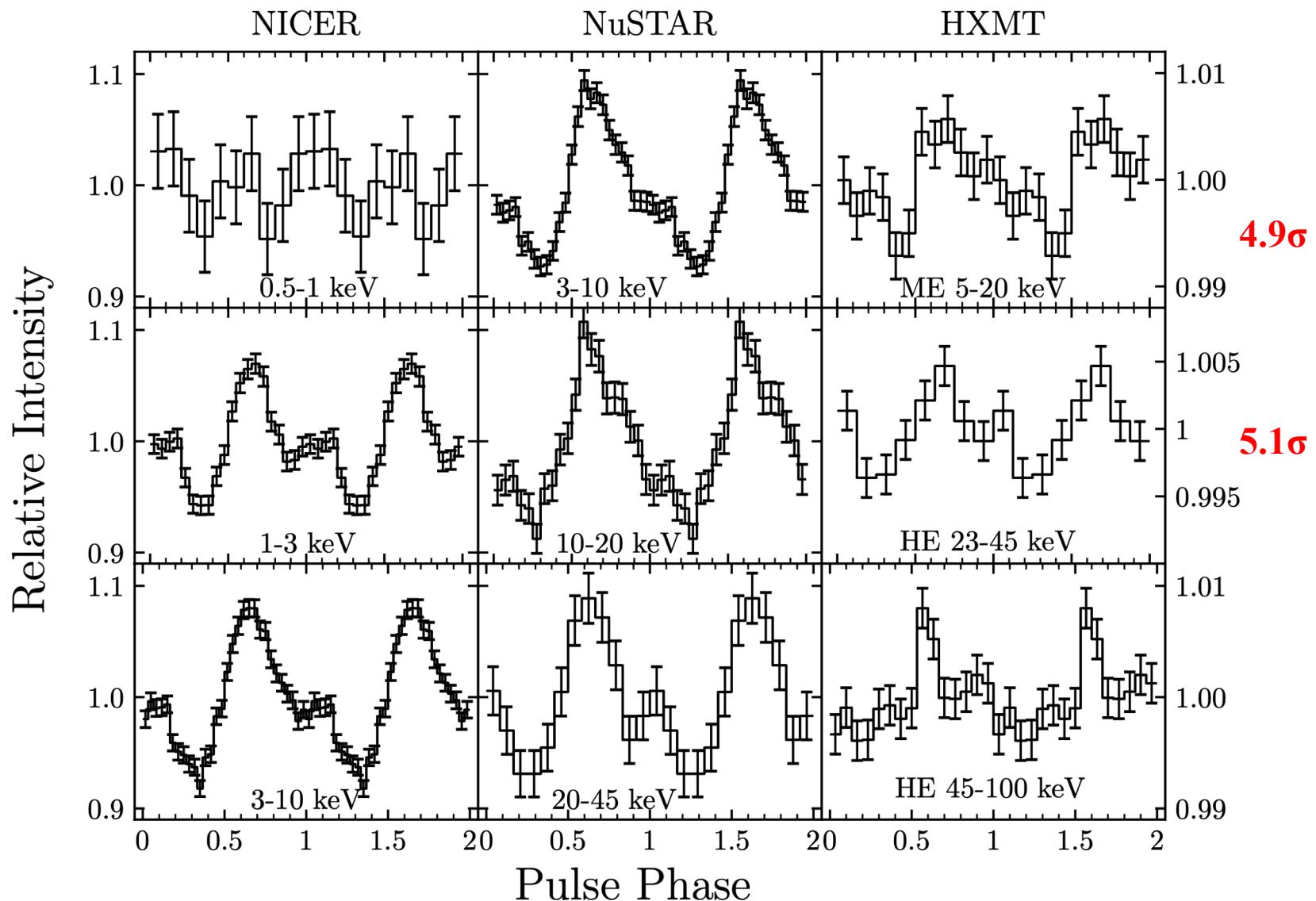
2018 & 2019 outbursts  
HXMT observed in 2018



**Table 3.** Positional, rotational, and orbital parameters used from other works and derived in this work for Swift J1756.9–2508.

Parameter	Units	Values
$\alpha_{2000}$		17 <sup>h</sup> 56 <sup>m</sup> 57 <sup>s</sup> 350
$\delta_{2000}$		-25°06'27.80
JPL Ephemeris		DE405
$P_{\text{orb}}$	s	3282.3515
$a_x \sin i$	lt-s	0.00597
$e$		0.00
Outburst – 2018		
$\nu$	Hz	182.065 803 84(3)
Epoch, $t_0$	MJD; TDB	58216.0
Validity range	MJD; TDB	58211–58219
$T_{\text{asc},2018}$	MJD; TDB	58211.017 52(6)
Outburst – 2019		
$\nu$	Hz	182.065 803 4(2)
Epoch, $t_0$	MJD; TDB	58654.0
Validity range	MJD; TDB	58653–58657
$T_{\text{asc},2019}$	MJD; TDB	58655.996 57(12)

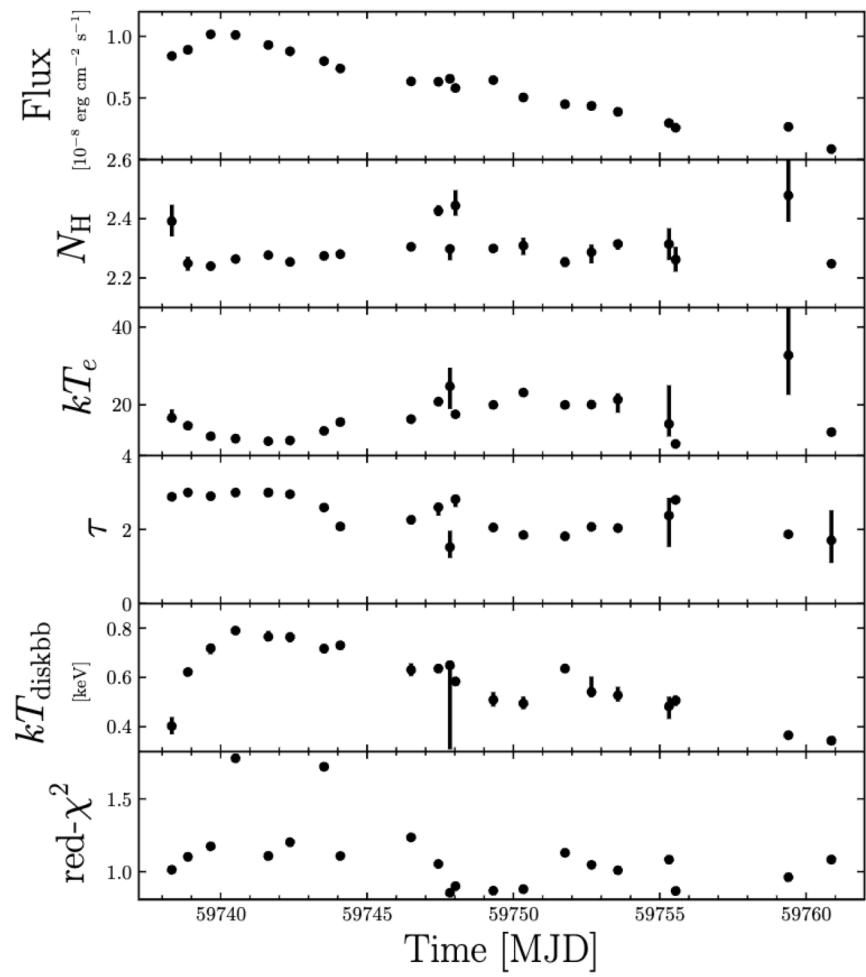
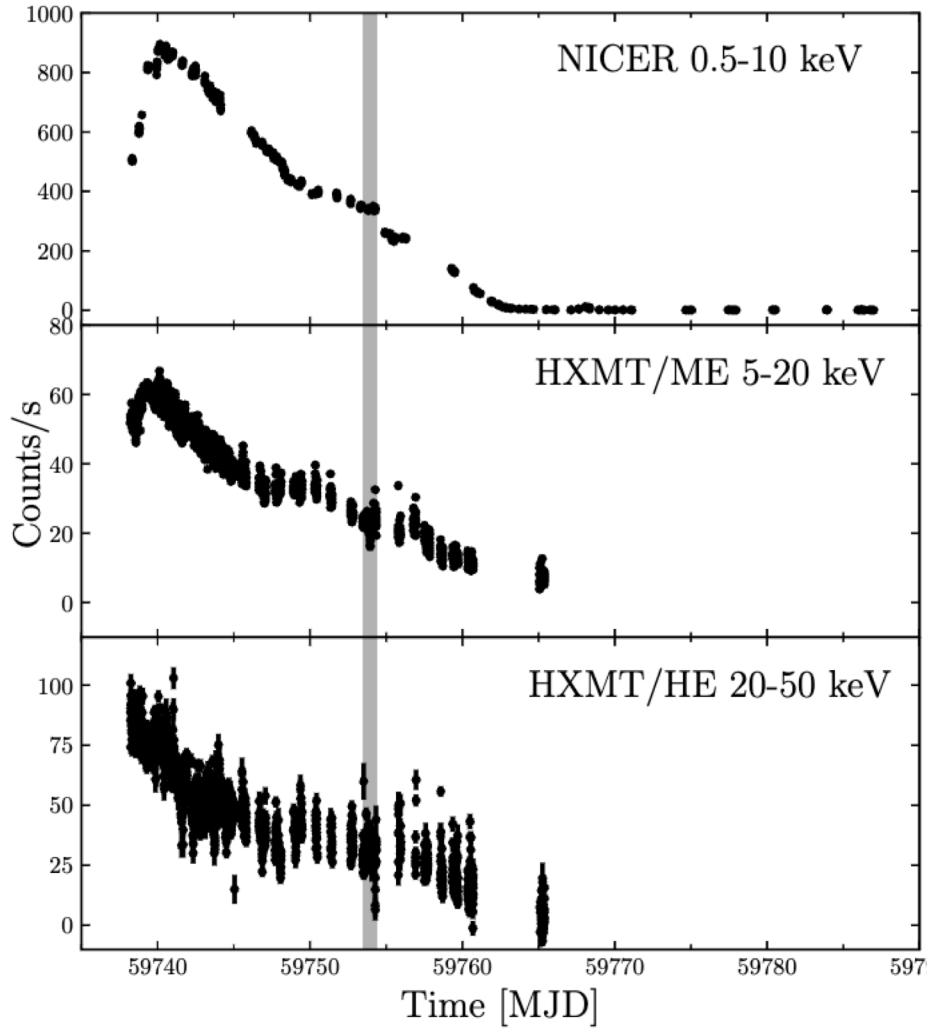
# Pulse profile-2018 outburst



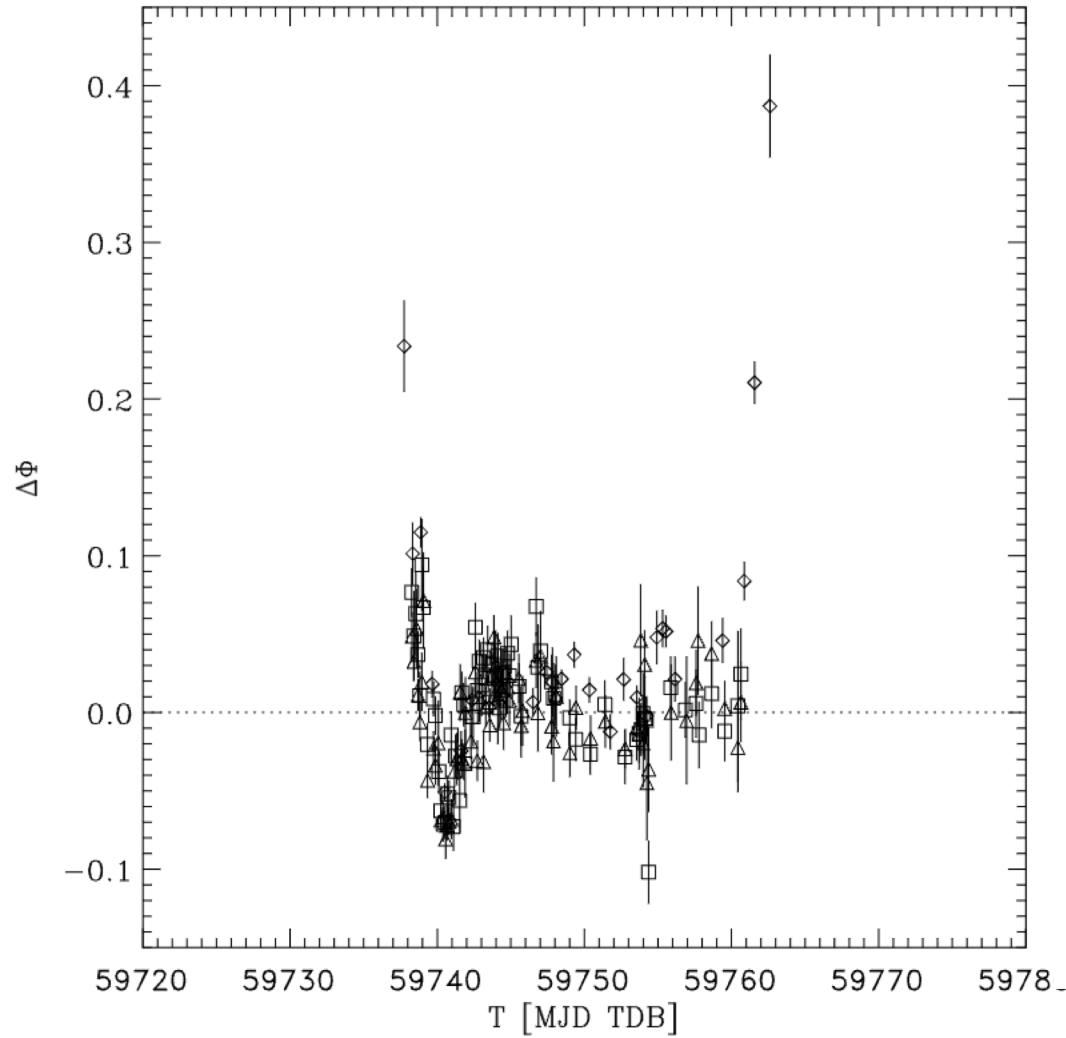
# Swift J1756.9-2508

During the 2018 (also for the 2019) outburst, no pulsed emission was detected below 1 keV using NICER and *XMM-Newton* data. *NuSTAR* detected pulsed emission up to  $\sim$ 60 keV during the 2018 outburst. Folding INTEGRAL-ISGRI barycentered data of the 2018 outburst for the 20–60 keV band did not result in a detection of the pulsed emission in line with the expectations given the moderate total outburst flux, low exposure, and low pulsed fraction of  $\lesssim$ 8% (see Sect. 5.2). However, with the *Insight-HXMT* HE instrument, significant pulsed emission has been detected up to  $\sim$ 100 keV (see lower right panel of Fig. 5) in spite of relatively low exposure. This suggests great prospects for future observations with *Insight-HXMT* of the AMXP outbursts.

# MAXI J1816-195 outburst



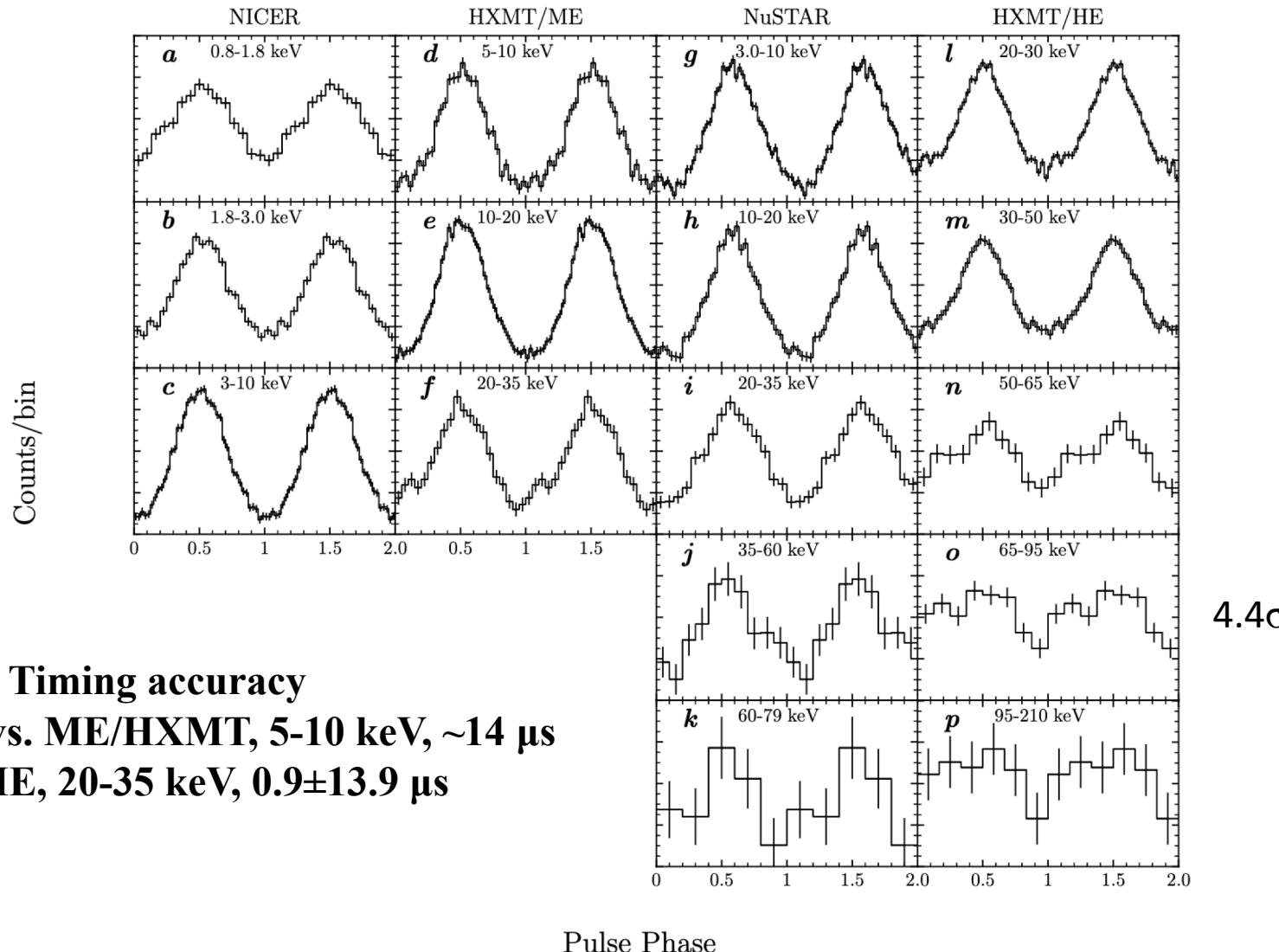
# MAXI J1816-195 timing



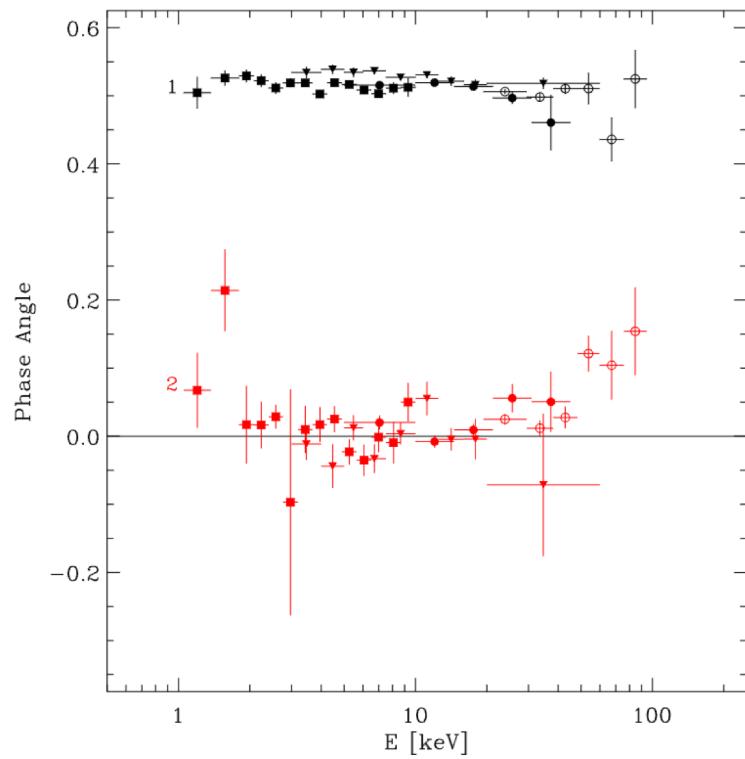
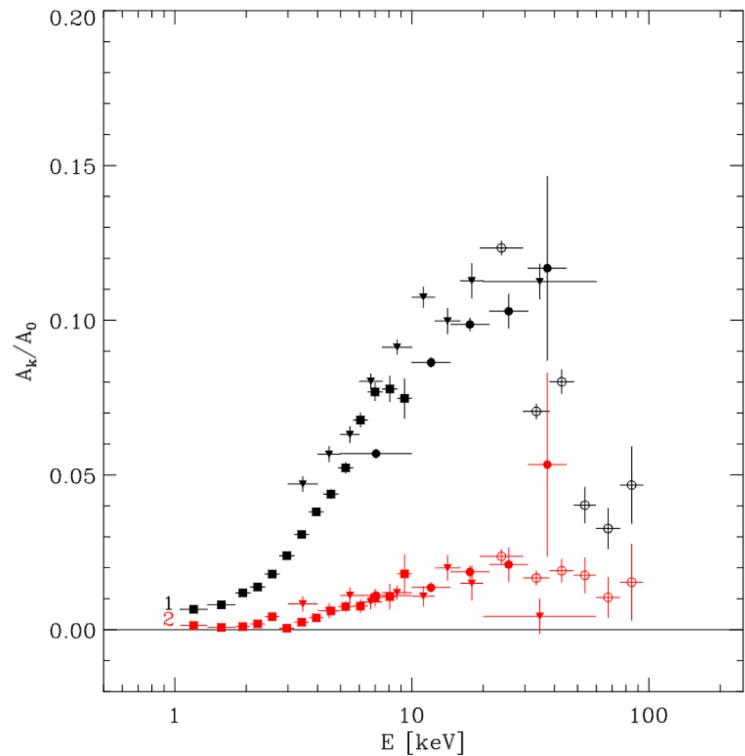
**Table 1.** Positional, orbital and spin parameters derived in this work and used as fixed values from literature for MAXI J1816–195.

Parameter	Values	Units
$\alpha_{2000}$	$18^{\mathrm{h}}16^{\mathrm{m}}52^{\mathrm{s}}41168(12)$	
$\delta_{2000}$	$-19^{\circ}37'57''40138(473)$	
JPL Ephemeris	DE405	
$P_{\mathrm{orb}}$	17402.5786(60)	s
$a_x \sin i$	0.262 948(18)	lt-s
$e$	0 (fixed)	
$T_{\mathrm{asc}}$	59738.875 632(4)	MJD (TDB)
Constant Frequency model		
Validity range	59741.9 – 59760.6	MJD (TDB)
$t_0$ (Epoch)	59741.0	MJD (TDB)
$\nu$	528.611 105 832(4)	Hz
$\chi^2/\mathrm{d.o.f}$	86.78/(47 – 1) = 1.886	
Spin-up model		
Validity range	59741.9 – 59760.6	MJD (TDB)
$t_0$ (Epoch)	59741.0	MJD (TDB)
$\nu$	528.611 105 774(12)	Hz
$\dot{\nu}$	$(9.0 \pm 2.1) \times 10^{-14}$	$\mathrm{Hz s}^{-1}$
$\chi^2/\mathrm{d.o.f}$	68.18/(47 – 2) = 1.515	

# MAXI J1816-195 pulse profile



# MAXI J1816-195 pulse profile



# Summary/Future

- **Swift J1756.9-2508**
  - ~20 ks (10 ks GTI) HXMT data produce  $\sim 5\sigma$  pulsation
- **MAXI J1816-195**
  - ~ 790 ks (180 ks GTI) HXMT data detected spin-up
- **Future works**
  - SAX J1808.4-3658, 2.8-3.5 $\sigma$  pulsation
  - IGR J17498-2921, ME (5-20 keV; 11.6  $\sigma$ ) and HE (20-60 keV; 18  $\sigma$ ; 60-90 keV  $\sim 3 \sigma$ )
  - Constraints on the NS mass and radius