



广州大学

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Multi-frequency Study On The Mode Switching Of PSR J0614+2229

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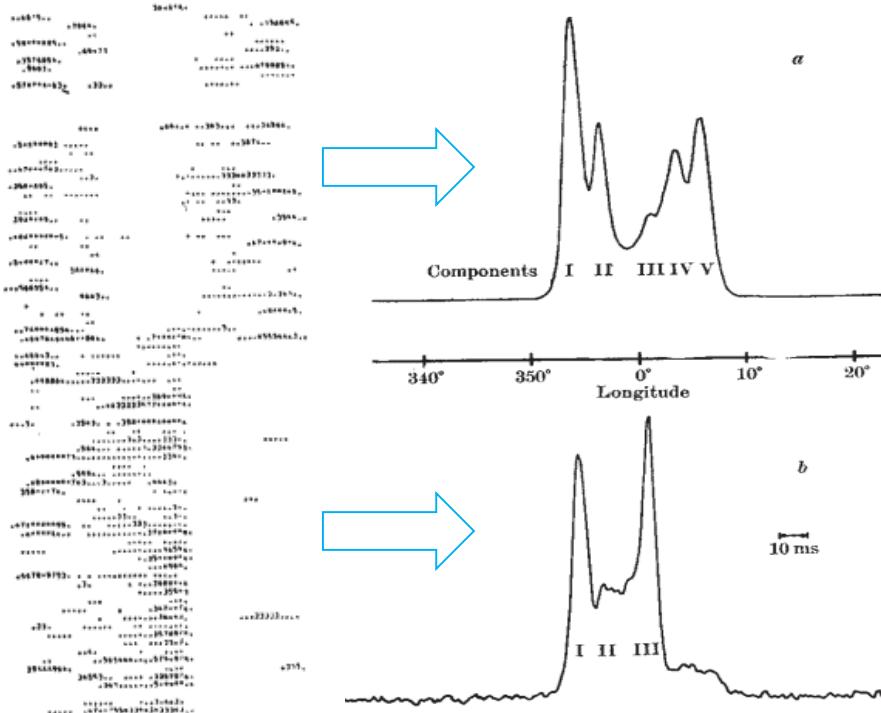
Outline

- ① Background
- ② Our Work
- ③ Summary
- ④ References

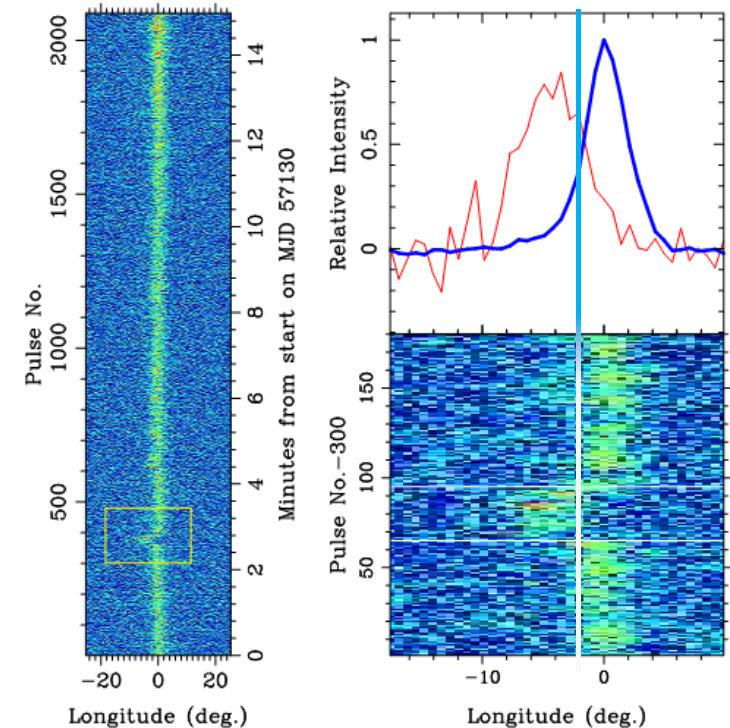
1. Background

- ◆ ~60 mode changing pulsars (hereafter MCPs) has been reported in 100+ papers (total no. of pulsars 2700+), ~5 MCPs with phase offset;

Phase stable

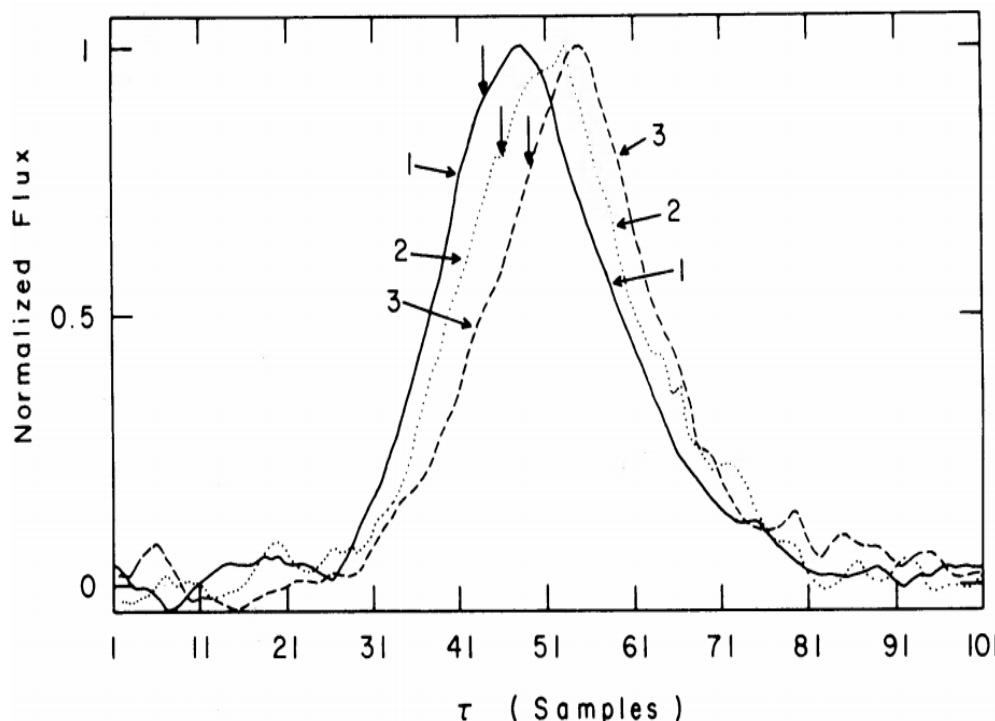


Phase offset

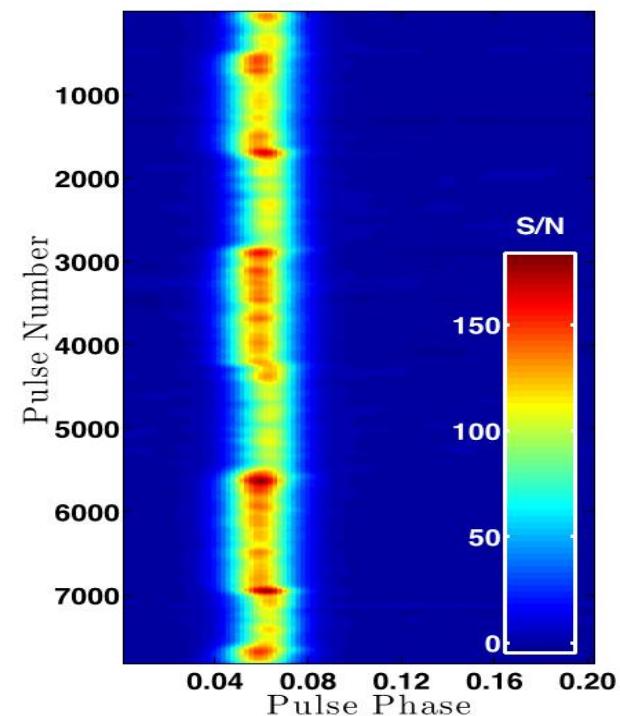


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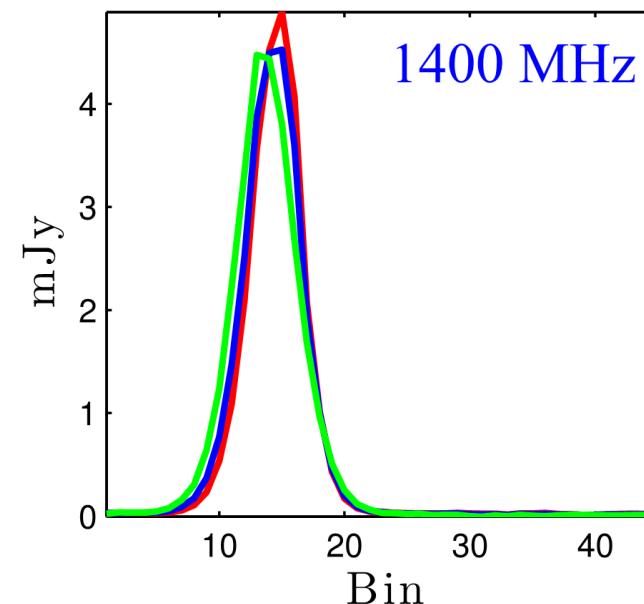
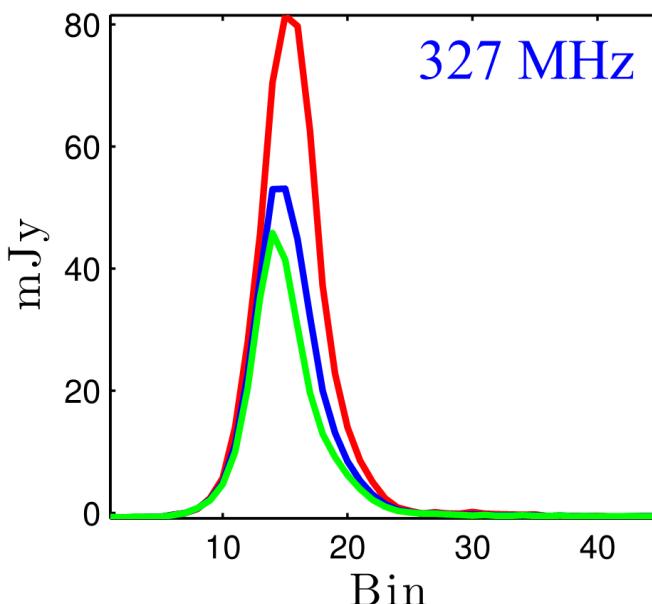
Ferguson & Boriakoff 1980



Rajwade et al. 2016

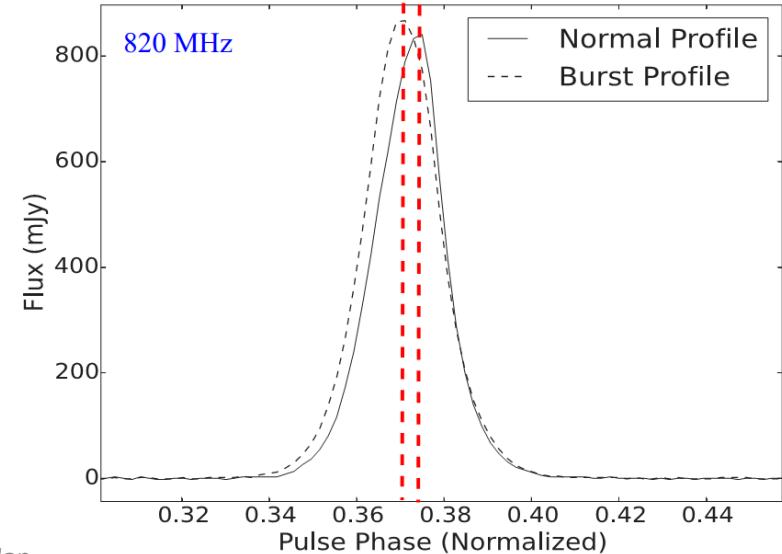
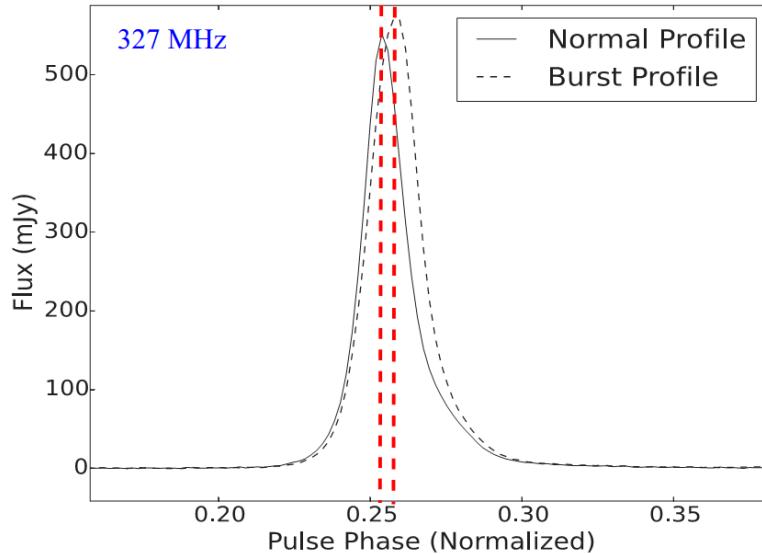
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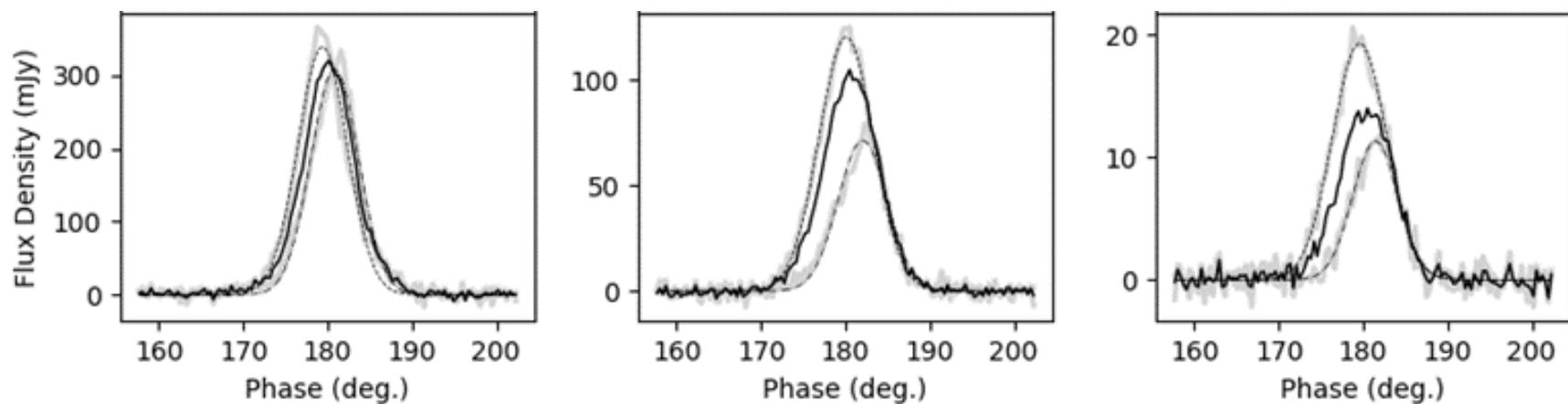
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- ◆ Seymour et al. (2014) reported that the phase of normal profile is more advanced than the burst state both at 0.3 GHz and 1.4 GHz in MJD 54898;
- ◆ Rajwade et al. (2016) noticed that the relationship between pulse phase and flux inverted at 0.8 GHz in MJD 56756 compared with 0.3/0.1 GHz.



2. Our work

- ◆ We processed the longest observation from Parkes' historical data in 50cm, 20cm and 10cm bands, and collected the data in the literature.

| ID | MJD (d) | Receiver | ν (MHz) | $\Delta\nu$ (MHz) | $\Delta\nu_{\text{ch}}$ (MHz) | T_{sub} (s) | N_{sub} | N_{bin} |
|-------|-------------|----------|----------------|----------------------|----------------------------------|-------------------------|------------------|------------------|
| 2005a | 53658.80926 | 5010CM | 686 | 256 | 0.125 | 59.6209 | 51 | 1024 |
| 2005b | 53667.80810 | H-OH | 1369 | 256 | 0.125 | 59.9563 | 18 | 1024 |
| 2005c | 53658.80926 | 1050CM | 3100 | 1024 | 1.0 | 59.6209 | 51 | 1024 |



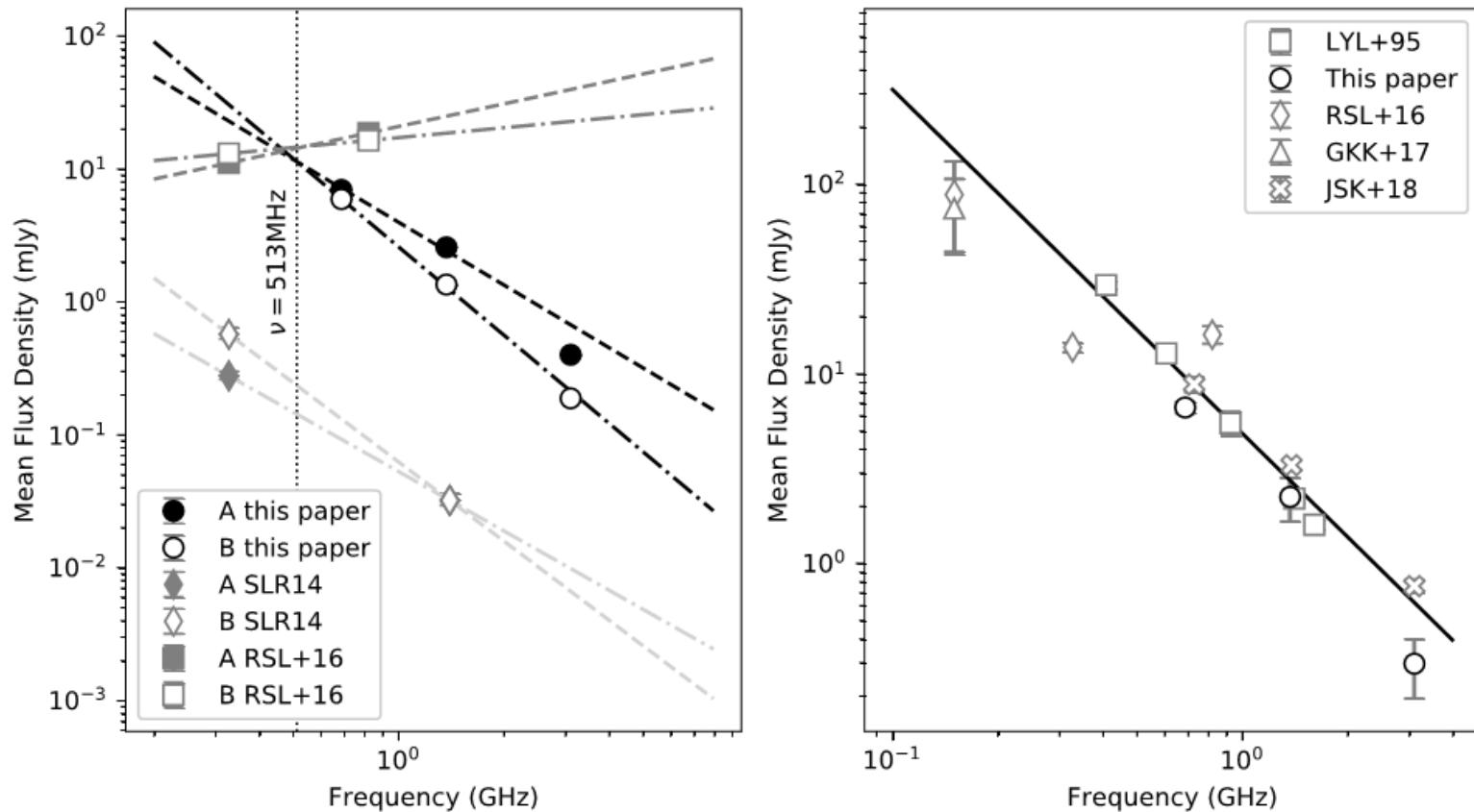
2. Our work

- ◆ Spectral difference between two modes.

| UT | MJD | Telescope | ν (MHz) | I_T (mJy) | I_A (mJy) | I_B (mJy) | I_A/I_B | Ref. |
|---------------|-------------|------------|----------------|----------------|-----------------------|-----------------------|----------------------|--------|
| | (d) | | | | | | | |
| 2005 Oct. 15 | 53658.80926 | Parkes | 653 | 6.7(5) | 7.0(4) | 6.0(2) | 1.2(1) | 2005a |
| 2005 Oct. 24 | 53667.80810 | Parkes | 1369 | 2.2(4) | 2.6(1) | 1.4(1) | 1.9(1) | 2005b |
| 2005 Oct. 15 | 53658.80926 | Parkes | 3100 | 0.3(2) | 0.40(5) | 0.19(3) | 2.1(4) | 2005c |
| 2009 Mar. 3-8 | 54893-54898 | Arecibo | 327 | — | 0.28(²) | 0.57(⁷) | 0.49(⁷) | SLR14 |
| 2009 Mar. 2-8 | 54892-54898 | Arecibo | 1400 | — | 0.032(⁴) | 0.032(⁴) | 1.0(²) | SLR14 |
| 2014 Apr. 9 | 56756 | LOFAR | 150 | 88(44) | — | — | — | RSL+16 |
| 2014 Apr. 9 | 56756 | Arecibo | 327 | 13.8(8) | 11.1(1) | 13.1(3) | 0.85(2) | RSL+16 |
| 2014 Apr. 9 | 56756 | Green Bank | 820 | 16.1(1.7) | 18.7(1.9) | 16.4(2) | 1.14(12) | RSL+16 |
| 2016 Jan. | 57391 | LOFAR | 150 | 75(32) | — | — | — | GKK+17 |
| 2007-2016 | — | Parkes | 728 | 8.8(8) | — | — | — | JSK+18 |
| 2007-2016 | — | Parkes | 1382 | 3.3(2) | — | — | — | JSK+18 |
| 2007-2016 | — | Parkes | 3100 | 0.76(6) | — | — | — | JSK+18 |

2. Our work

- ◆ Spectral difference between two modes.



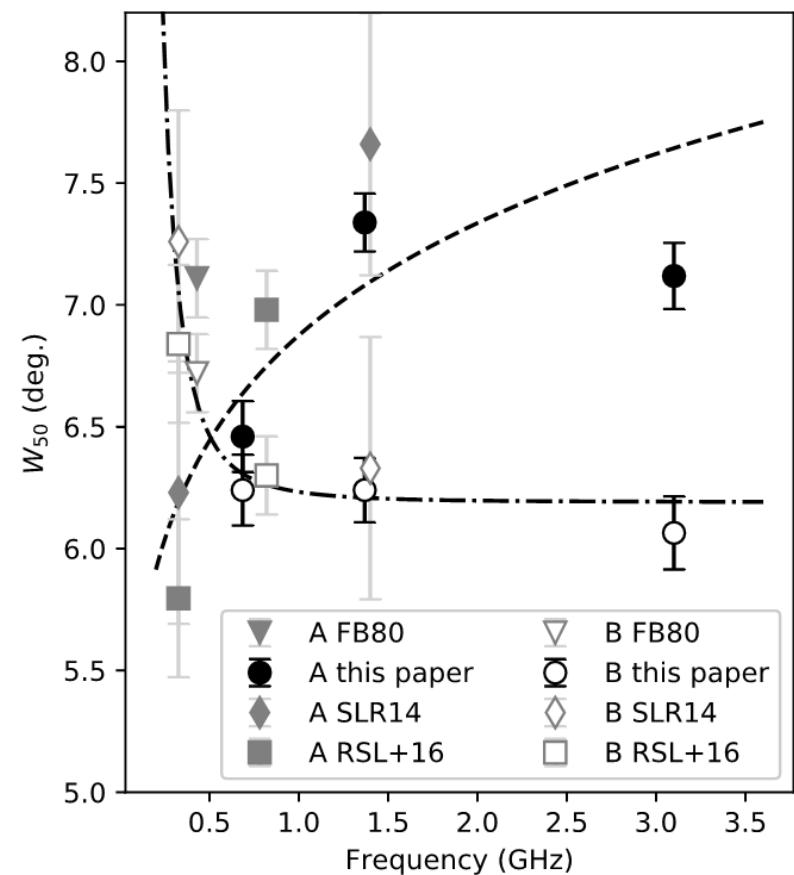
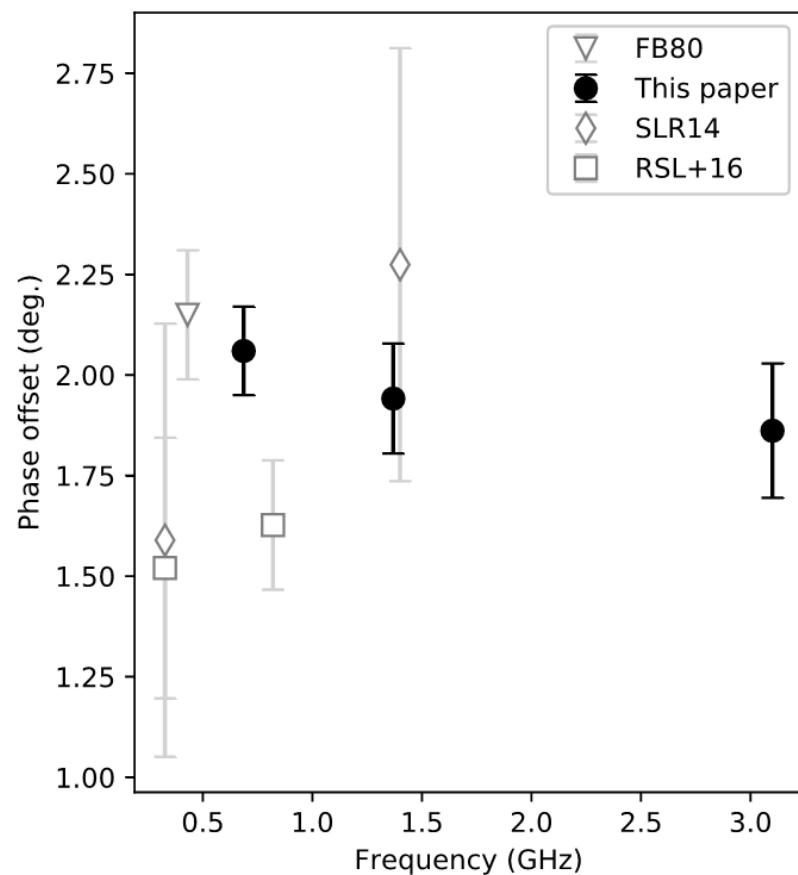
2. Our work

- ◆ Frequency dependence of phase offset and pulse width.

| ν (MHz) | $\Delta\Phi$ (deg.) | $\Delta\mu$ (deg.) | $W_{50,A}$ (deg.) | $W_{50,B}$ (deg.) | Ref. |
|----------------|------------------------|-----------------------|----------------------|----------------------|--------|
| 686 | 2.81(16) | 1.44(14) | 9.46(15) | 6.24(15) | 2005a |
| 1369 | – | 1.94(13) | 7.34(12) | 6.24(13) | 2005b |
| 3100 | – | 1.86(17) | 7.12(14) | 6.06(15) | 2005c |
| 430 | 2.15(16) | – | 5.15(16) | 4.73(16) | FB80 |
| 327 | 1.59(54) | – | 6.23(54) | 7.26(54) | SLR14 |
| 1400 | 2.27(54) | – | 7.66(54) | 6.33(54) | SLR14 |
| 327 | 1.52(32) | – | 5.80(32) | 6.84(32) | RSL+16 |
| 820 | 1.63(16) | – | 6.98(16) | 6.30(16) | RSL+16 |

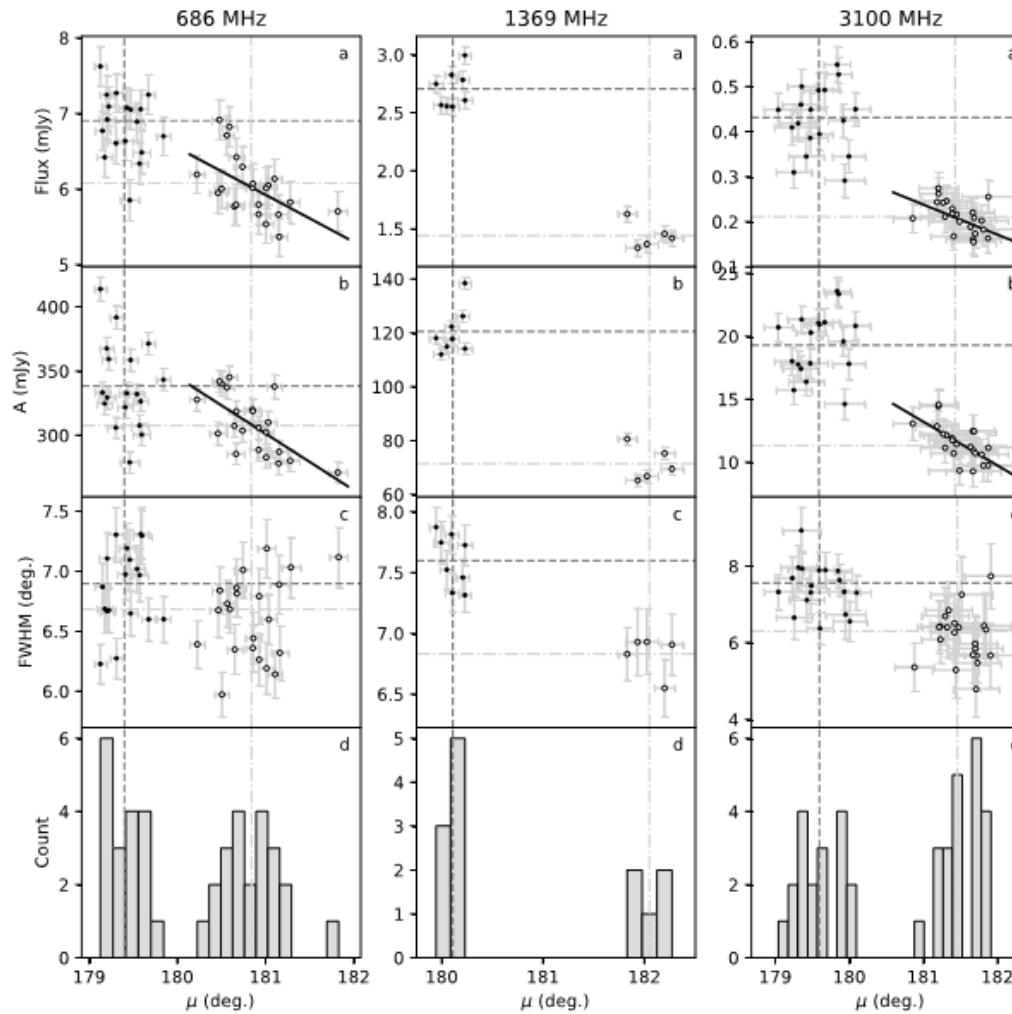
2. Our work

- Frequency dependence of phase offset and pulse width.



2. Our work

- ◆ Correlation between the peak amplitude and phase for mode B.



2. Our work

- ◆ 脉冲辐射的偏振位置角(PPA)可以由RVM模型描述；
- ◆ 经典RVM模型和SGP点；

$$\tan(\psi - \psi_0) = \frac{\sin \alpha \sin(\phi - \phi_0)}{\sin \zeta \cos \alpha - \cos \zeta \sin \alpha \cos(\phi - \phi_0)}$$

$$\left(\frac{d\psi}{d\phi}\right)_{\max} = \left(\frac{d\psi}{d\phi}\right)_{\phi_0} = \frac{\sin \alpha}{\sin \beta}$$

- ◆ 相对论性RVM模型；

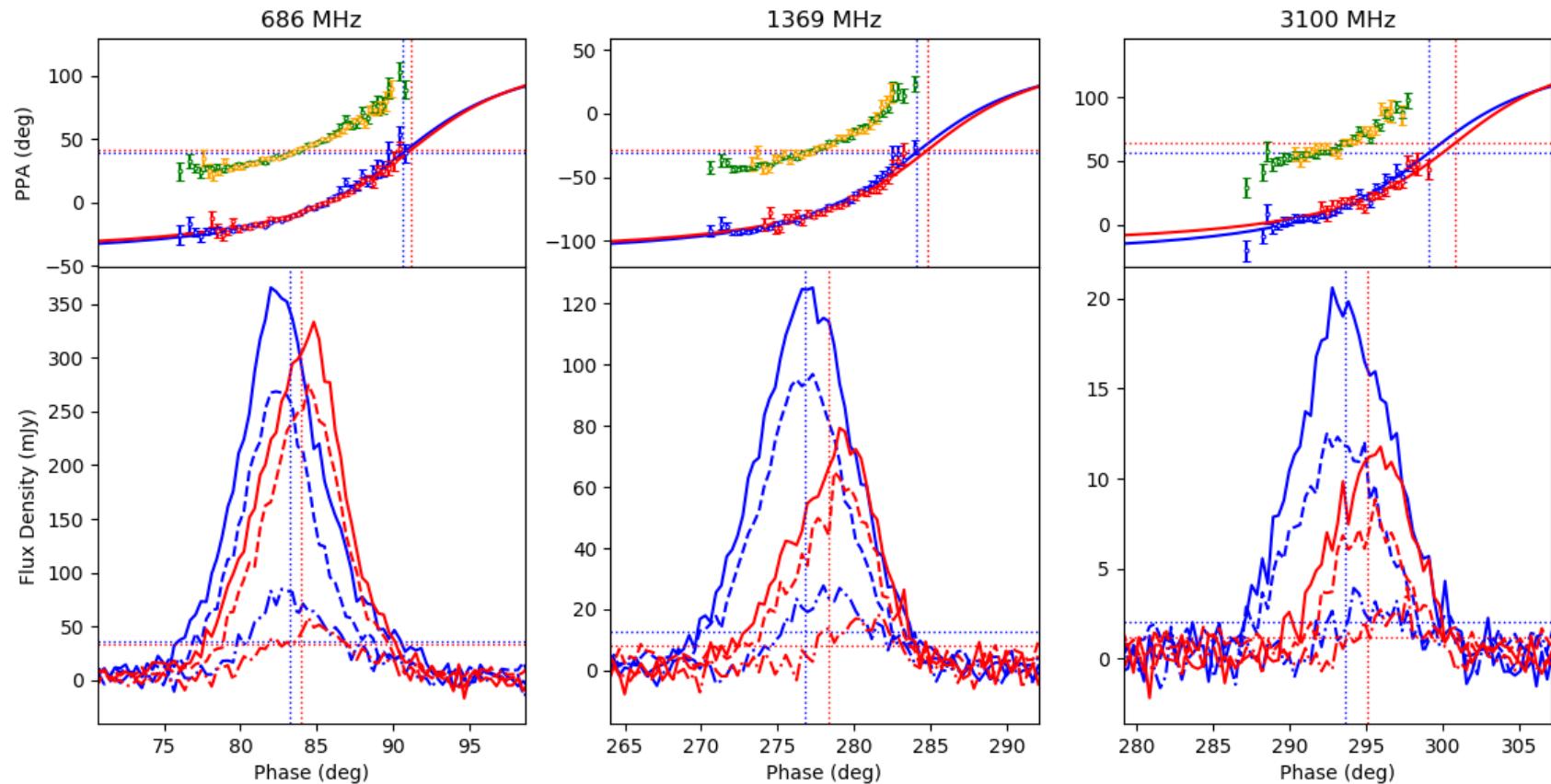
$$\tan(\psi_r - \psi_\omega + \Delta\psi_J) \approx \frac{\sin \alpha \sin(\phi_{\text{obs}} - \phi_f + \Delta\phi_r)}{\sin \zeta \cos \alpha - \cos \zeta \sin \alpha \cos(\phi_{\text{obs}} - \phi_f + \Delta\phi_r)}$$

- ◆ 辐射高度。

$$\Delta\phi_r \approx -\frac{2r}{R_{lc}}$$

2. Our work

◆ 偏振



2. Our work

- ◆ 辐射区域
- ◆ 辐射区高度变化: $-70 \pm 40 \text{ km}$ 、 $-110 \pm 70 \text{ km}$ 、 $-270 \pm 150 \text{ km}$

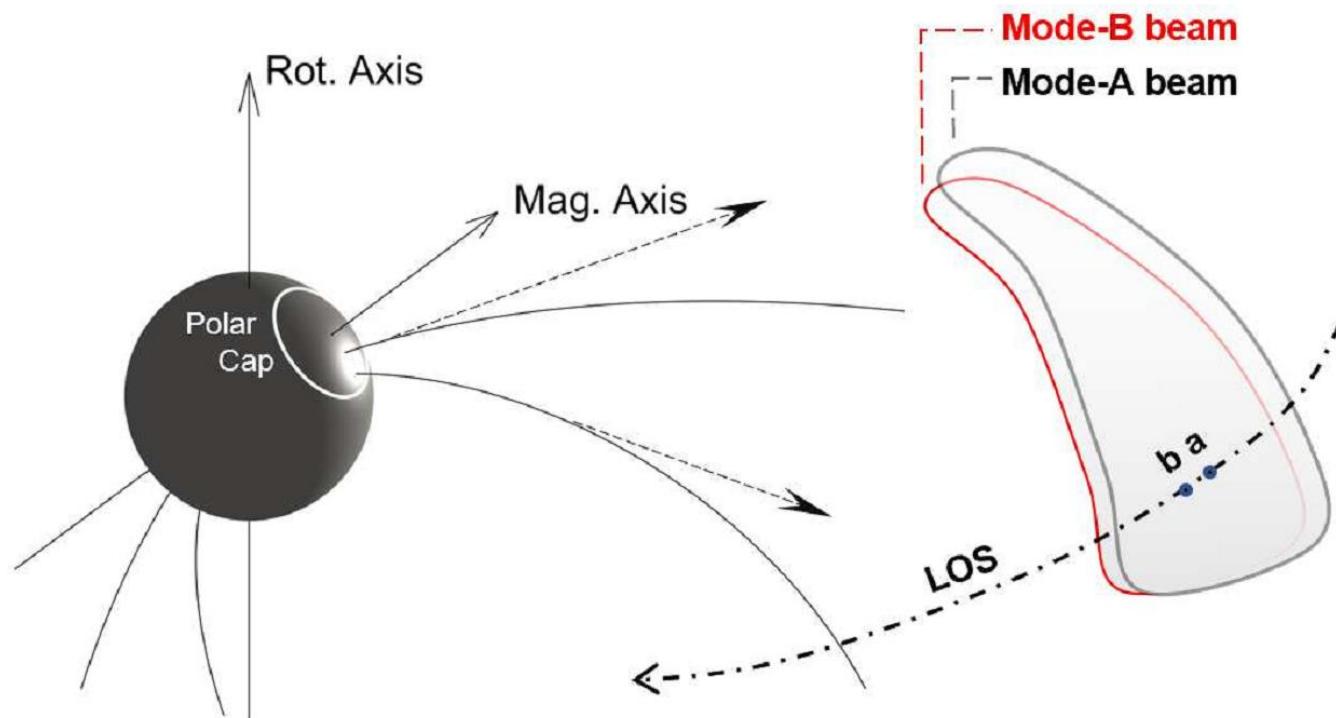
表 4.1 脉冲星 PSR J0614+2229 辐射窗口参数

| Freq/MHz | Mode | $L_{10}/^\circ$ | $R_{10}/^\circ$ | $\phi_{\text{prof}}/^\circ$ | $\phi_0/^\circ$ | $\psi_0/^\circ$ | r_{BCW}/km |
|----------|------|-----------------|-----------------|-----------------------------|-----------------|-----------------|----------------------------|
| 686 | A | 76.0(6) | 90.5(1.1) | 83.3(6) | 90.7(2) | 39.3(9) | 520(40) |
| 686 | B | 77.6(5) | 90.2(5) | 83.9(4) | 91.2(2) | 41.7(9) | 510(30) |
| 1369 | A | 269.8(7) | 283.8(5) | 276.8(4) | 284.1(2) | -30.8(9) | 510(30) |
| 1369 | B | 271(2) | 284(1) | 278(1) | 284.9(5) | -28(2) | 480(80) |
| 3100 | A | 287(1) | 300.5(8) | 293.8(7) | 299.1(5) | 57(2) | 370(60) |
| 3100 | B | 289(2) | 300.3(9) | 294.6(9) | 301(1) | 64(6) | 450(90) |

注: 括号中为参数末位的误差值。

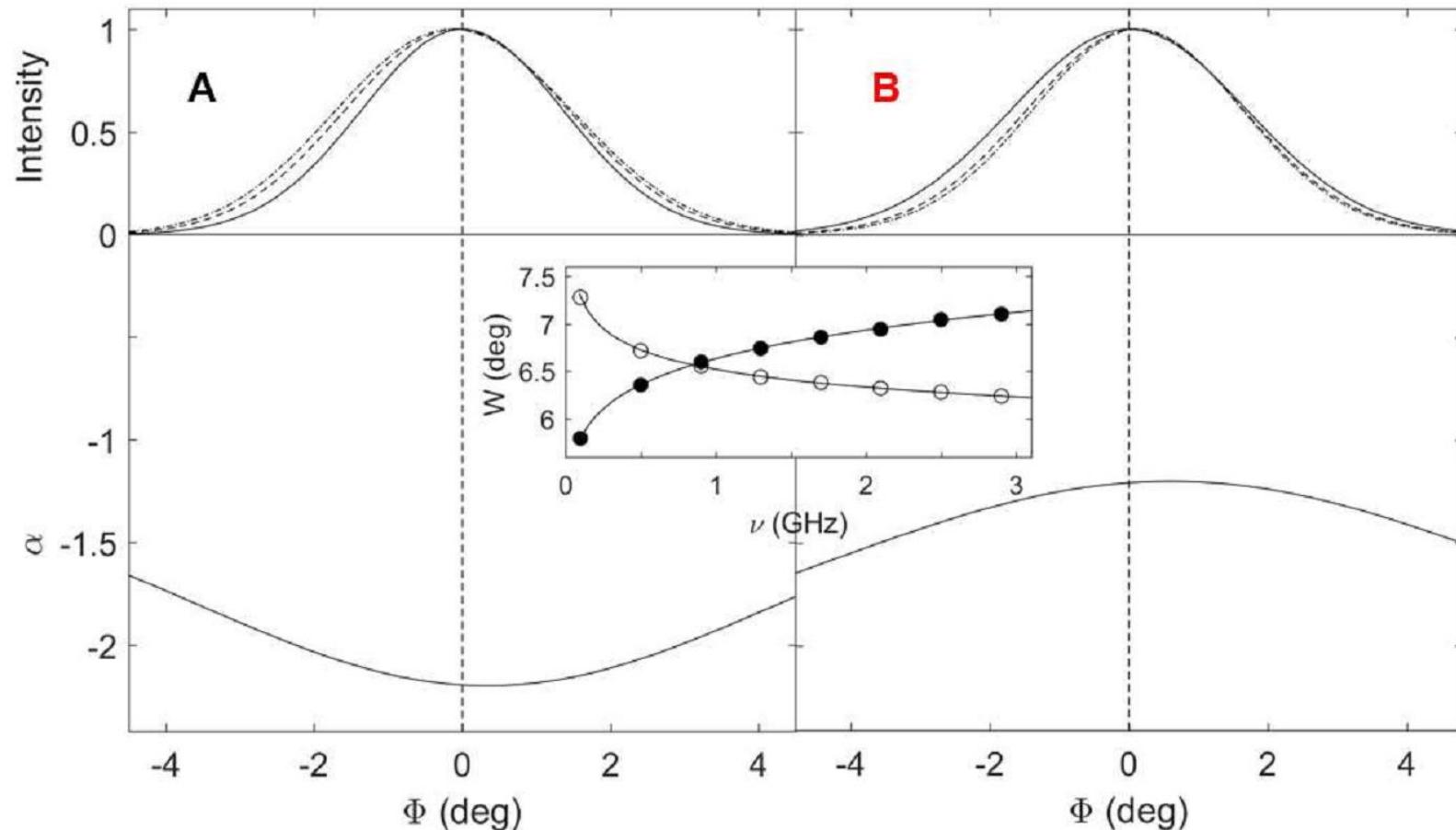
2. Our work

- ◆ 辐射区域
- ◆ 辐射区高度变化: $-70 \pm 40 \text{ km}$ 、 $-110 \pm 70 \text{ km}$ 、 $-270 \pm 150 \text{ km}$
- ◆ 辐射区经度变化:



2. Our work

- ◆ Implication to spectral properties of the emission beams.



3. Summary

- ◆ Mode dependent spectra
- ◆ Mode dependent Thorsett relationships
- ◆ Frequency independent phase offset
- ◆ Mode dependent phase-flux correlations
- ◆ 磁场位形在模式变换时基本不变
- ◆ SGP点偏移方向和辐射窗口中心偏移方向相同

- ◆ 辐射区高度: $A < B$
- ◆ 辐射区经度: $A < B$
- ◆ Different spectral distributions

- ◆ Phase-resolved spectra

4. References

- [1] 张颜荣,陈建玲,王洪光,黄秀健,2019. 模式变换脉冲星PSR J0614+2229的多波段辐射区的研究[J]. 天文研究与技术.
- [2] 张颜荣,王洪光,黄秀健,陈建玲,2019. Multifrequency Study on The Mode Switching of PSR J0614+2229[J]. ApJ. (Under review)
- [3] 张颜荣. 脉冲星模式变换的分类研究[D]. 广州大学, 2019.

Thanks for Your Attention!