

European E Commission

Horizon 2020 European Union funding for Research & Innovation

ARG

Conceptual Design Study

Designing a Next-Generation Radio Telescope for Multi-Messenger Astronomy

Introduction to Argos

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DDF2024 May 10 – 15, 2024





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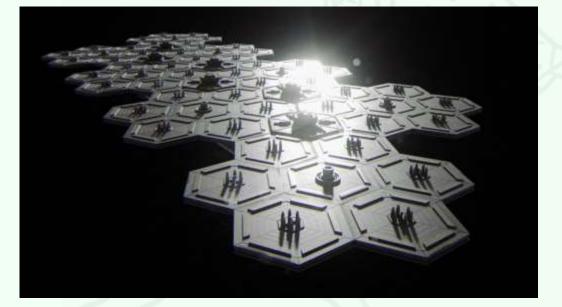






['dːgɒs] Ἄργος Πανόπτης, "All-seeing Argos" : A many-eyed giant in Greek Mythology, the guardian of the heifernymph lo and the son of Arestor.

...a popular name for telescopes



Argos Subspace Radio Telescope Array in service with Starfleet since 2259



Argos array picked up techno-signatures in Carl Sagan's "Contact" (played by the VLA in the movie)

ARray for Gigahertz ObservationS

A telescope that combines broad frequency coverage, wide field, high sensitivity, high time and spatial resolution at GHz frequencies

...at a low price

Also: a facility that produces science-ready data with low lattency

Why bother? From discoveries to high-impact science

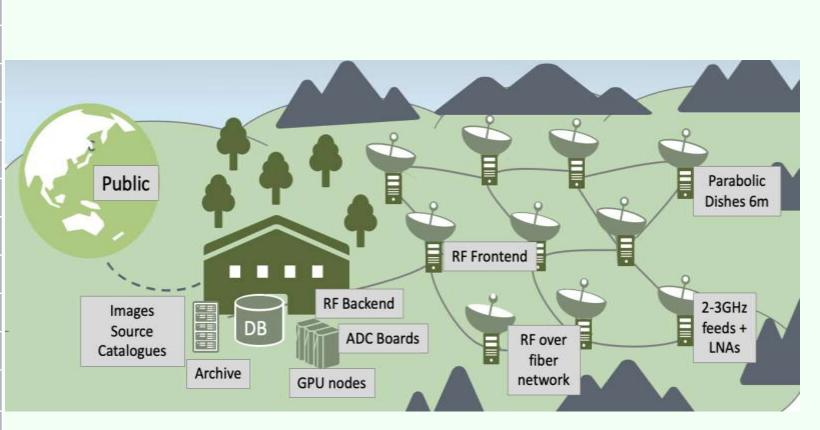


Why bother? Panoptic surveys are revolutionizing astronomy Credit: NASA/Fermi

ARGOS system overview



antennas	~1200 x 6m
collecting area	33,000 m ²
instantaneous FoV	~9.6 deg ²
system temperature	Ambient LNAs (<35K)
apperture efficiency	>70%
resolution	5"
tracking	EL/AZ
frequency	1—3 GHz
SEFD	6.3 Jy
rms noise/beam [30 min]	2.3 µJy
pulsar sensitivity [10 min]	0.02 mJy
survey speed figure-of-merit	$10^8 \text{ deg}^2 \text{ m}^4 \text{ MHz} / \text{K}^2$
FRB localisation accuracy	O(1")
Cost (construction)	<50,000 €/dish — 60 M€



ARGOS system overview



Why this design?...main advantages

- Low-cost
- Scalability
- Expansibility / Modularity
- Complementarity to other facilities

- → 60M EUR construction cost
- \rightarrow 50K EUR / element for full array (please adopt an antenna)
- ightarrow Science modules subscribe to same data stream
- → Wide-field/High-res; freq coverage, long/lat. EU priorities
- Incentives for private-sector involvement → COTS components / tech with spill-over applications
- Legacy value

→ Science-ready data

The ARGOS Science Themes

ARG

The Violent Universe Multi-messenger



The Evolving Universe Pulsar Timing Arrays

Next-generation telescopes Technology exploitation



Fundamental Physics Pulsars

> **Cosmology** Fast Radio Bursts

The Dynamic Universe Imaging



Fast Radio Bursts



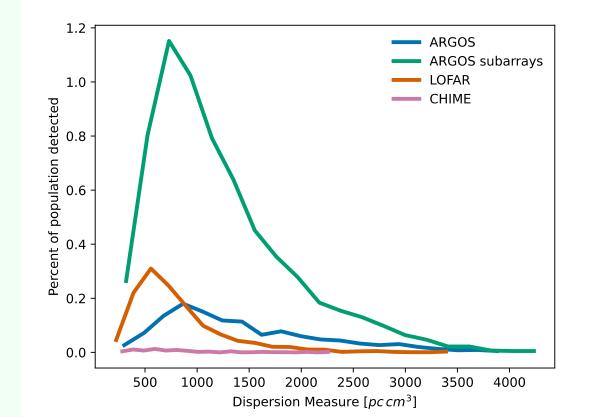
ARGOS will be a superb FRB discovery engine

Key specifications

10 deg² field of view, 1—3 GHz, 4km baselines, SKA-MID1 level collecting area, 250kHz channels, 30µsec time resolution, default search mode for dispersion measures up to 3000 pc cm⁻³, sub-arraying, low-lattency alerts

- Commensal mode
- One well-localised (<5") FRB per hour*</p>
- Sensitive to z>3 FRBs
- Flux, polarization, DM, RM information
- Automated low-latency triggers

Rate simulation (T2.2) assuming 70% uptime for ARGOS core, 20% channel corruption due to RFI *5 - 10 unique sources per day; 20 - 60 bursts

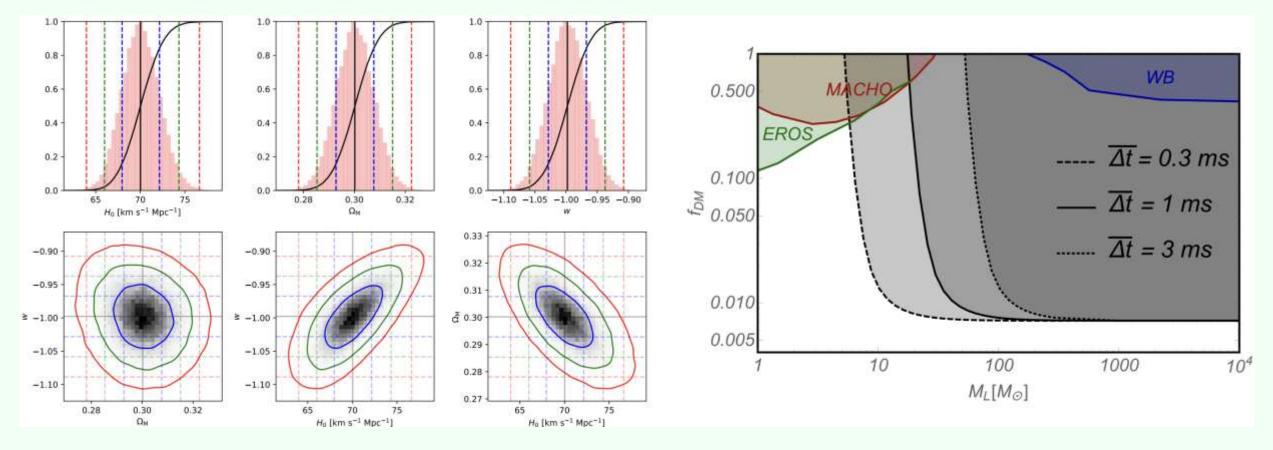


Fast Radio Bursts



Ultra-high precision cosmology with FRBs

2-5 strongly-lensed FRBs per year; will enable high precision cosmology and dark matter searches



Wuckniz et al. (2020)

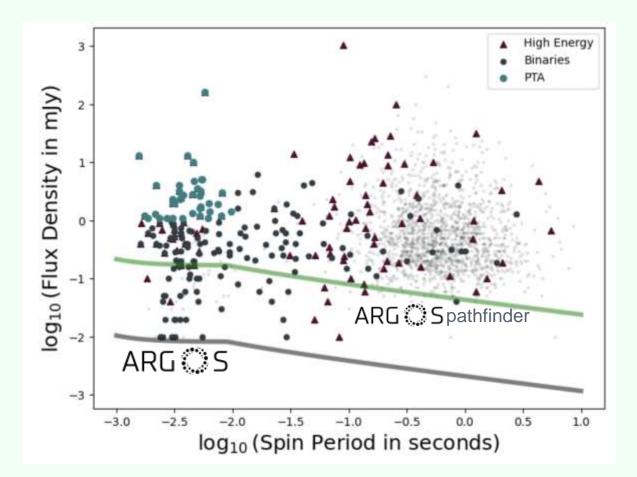
Munoz et al. (2016)

Pulsars

Key specifications

Up to 16 phased beams, 1—3 GHz, 4km baselines, SKA-MID1 level collecting area, 20kHz channels, 1µsec time resolution, coherent de-dispersion, sub-arraying, dynamic spectra, automated TOA generation

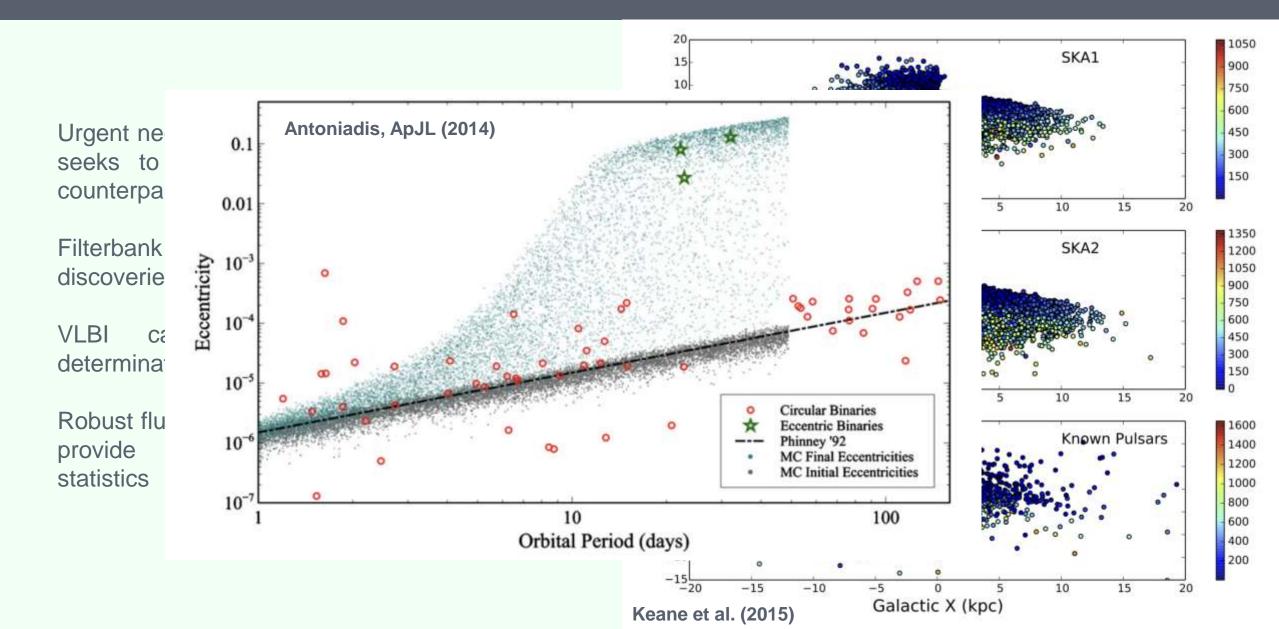
- approximately 1 pulsar per square degree on average → several 100s of pulsars per day;
- Coherent de-dispersion of known sources to maximize SNR
- Sub-banded times of arrival
- Flux, polarization information, DM, RM information





Pulsars: SKA/FAST follow-up

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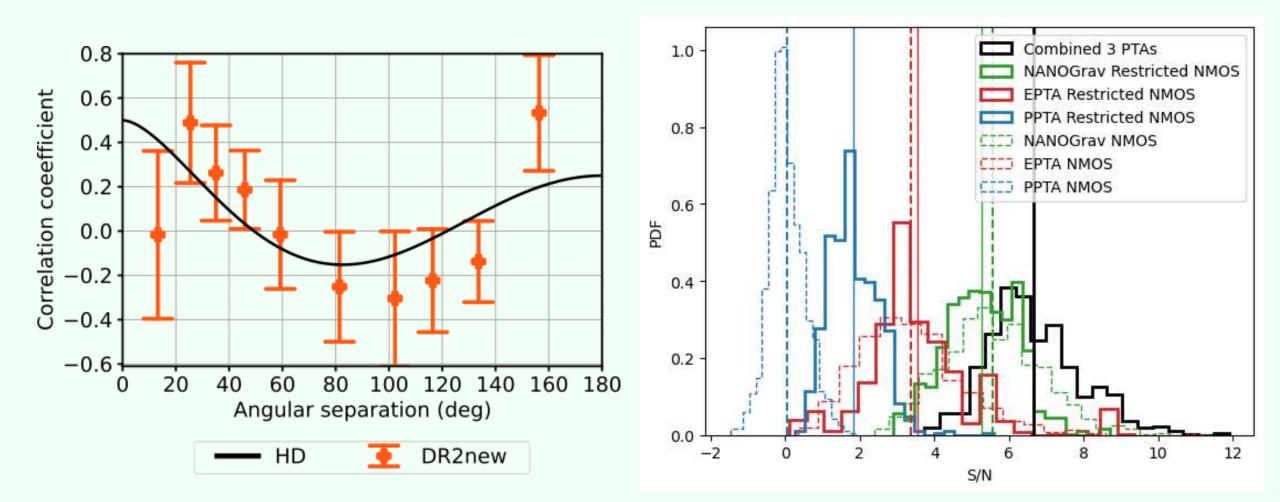
Pulsar timing arrays





The EPTA DR2 – A new GW window

Signal consistent with a nanohertz stochastic GW background. Also detected by NANOGrav

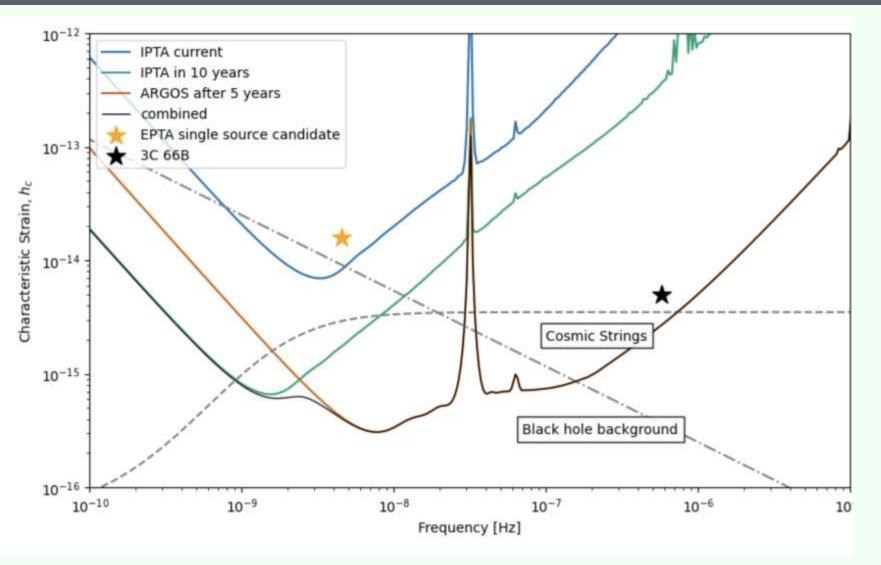


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EPTA, InPTA, IPTA (Antoniadis et al. 2023a-e, Smarra et al 2023, Agazie et al. 2024)

PTA science with ARGOS





ARGOS will improve PTA sensitivity in three ways

- ✓ Broadband TOAs + monitoring of profile changes and ISM will enable effective removal of timing noise
- ✓ Large number (>10⁶) of high SNR TOAs
- ✓ High cadence will extend sensitivity to µHz frequencies. This is critical as backgrounds of different origins are expected to differ most at higher frequencies.

Also, key requirement for multimessenger astrophysics.

Projected sensitivity assuming 70% uptime, 20% channel corruption due to RFI, 30 PSRs, SNR=1000

Imaging and image-plane transients

ARG

 Quasar IOPTICAL Set Quasar (RADIO Sel. Stellar V CV

Line: $N_{exp} = 1$

8 Radio Pulsar X-ray binary

GRB ▲ Supernova

10

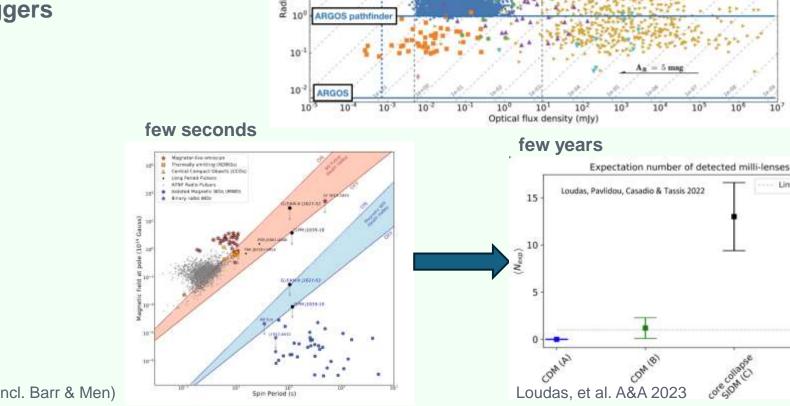
Optical AB magnitude (mag)

14

12

10

- High-cadence, full-Stokes imaging of 500 deg² per day
- Light curves and transient alerts
- Follow-up of multi-messenger triggers
- Full-stokes images, visibilities, source catalogues, classifications, triggers



Credit: R. Fender

105

10

10 (M/m) 10 Sitv

Xnu 10 24

22

20

18

16

Rea et al. (2023) ApJL; Hurley-Walker et al., 2023, Nature (incl. Barr & Men)

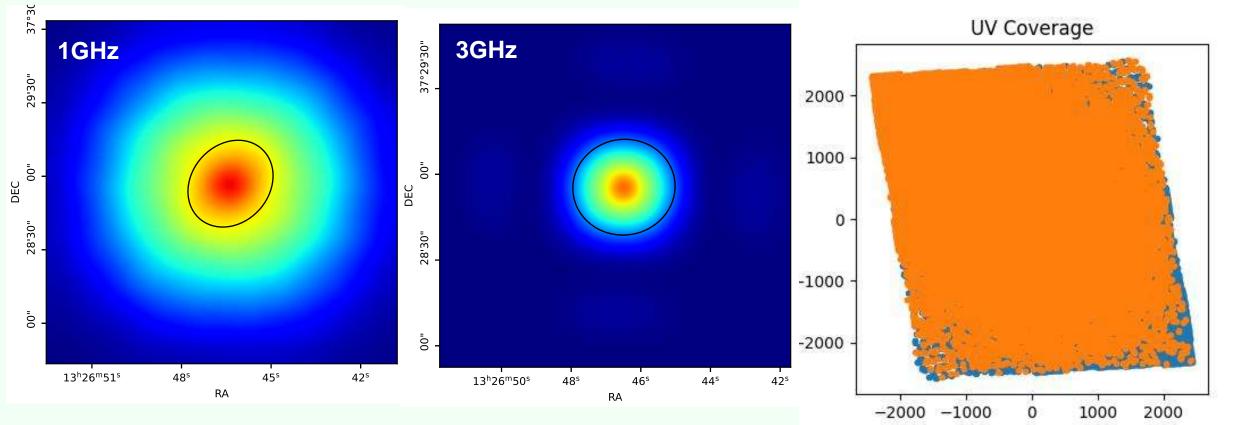
Imaging and image-plane transients



ARGOS imaging

Key specifications

Regular 32x32 grid (<530m) + gaussian random (< 4 km) configuration, AI-based flagging and classification



Snapshot (5min) UV coverage and dirty beam profiles for a dec=+40° source (T2.2) https://github.com/wchenastro/Mosaic

The ARGOS Conceptual Design Study

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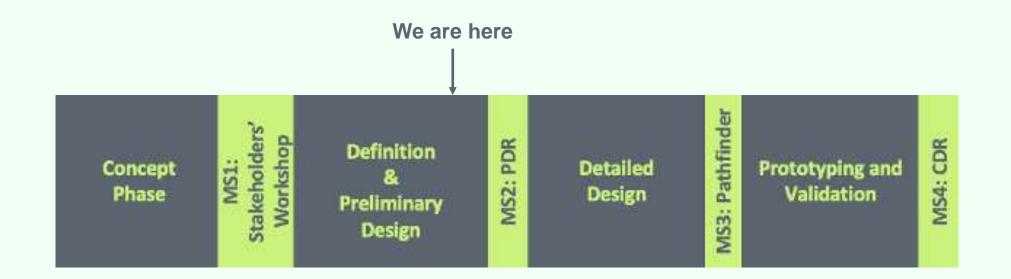


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This project has received funding from the European Union's Horizon Europe research and innovation programme under grant agreement **No 101094354**. The content of this material reflects the opinions of its authors and does not in any way represent the opinions of the European Union

Timeline

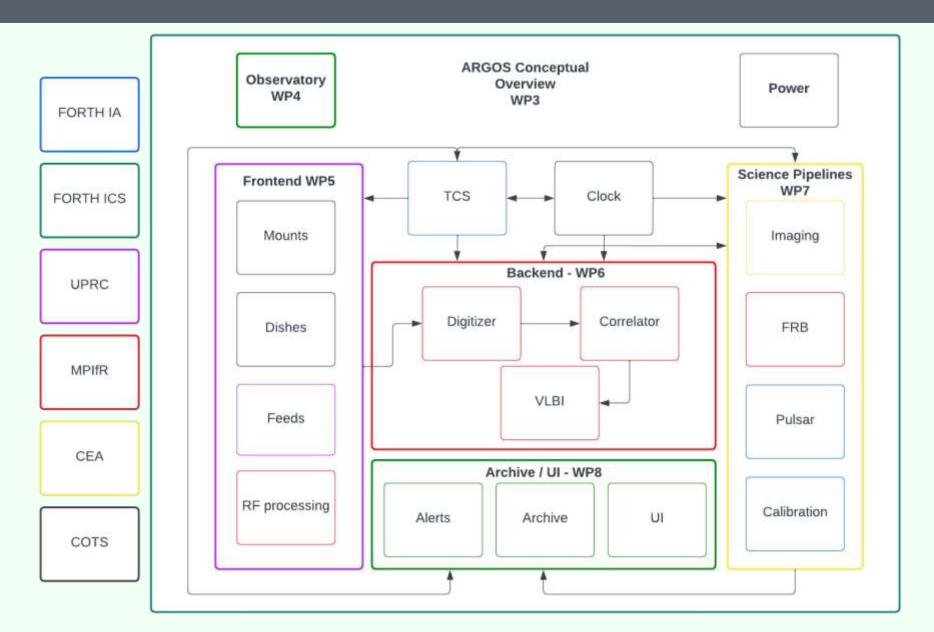




Final Deliverables:

10-dish prototype detailed fully-costed design Structure





Frontend and Digital Feed

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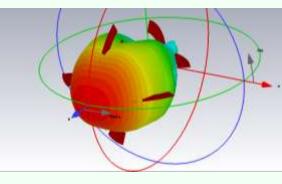
Mechatron SOLARIBACKER



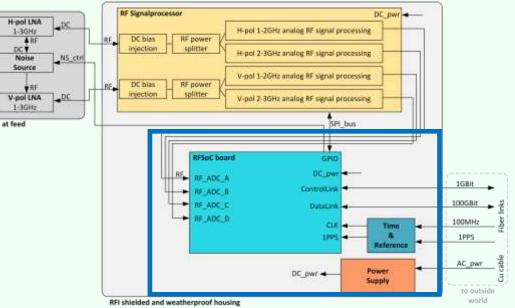
UNIVERSITY OF PIRAEUS RESEARCH CENTER

Max-Planck-Institut für Radioastronomie

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Low-cost mass fabrication of the feed system ARG



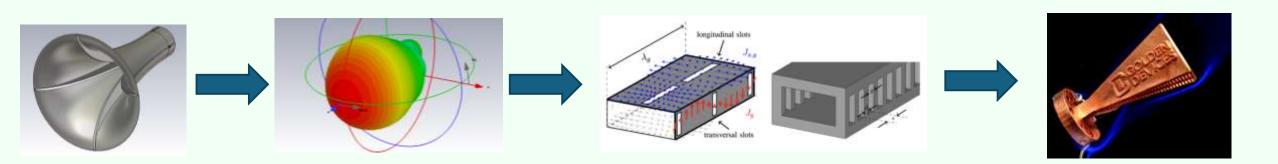
SLS (Selective Laser Sintering) 3D-printing

(a) lightweight structure (3kg)(b) reduced production time(c) low fabrication cost(d) durability

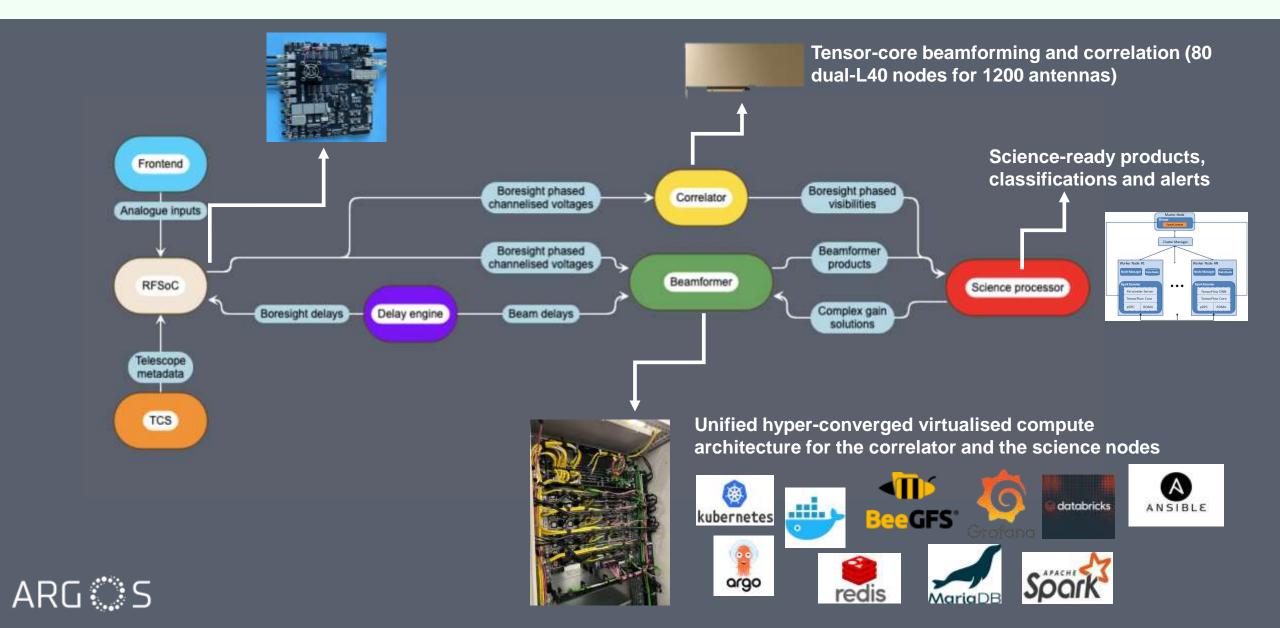
Key Innovations

- Monolithic structures with non-radiating slots
- Nickel coating, 7-50µ
- Plastic coating and front cover (EM transparent; FDM ASA ($\varepsilon r = 2.8$, tan $\delta = 7 \times 10^{-3}$)

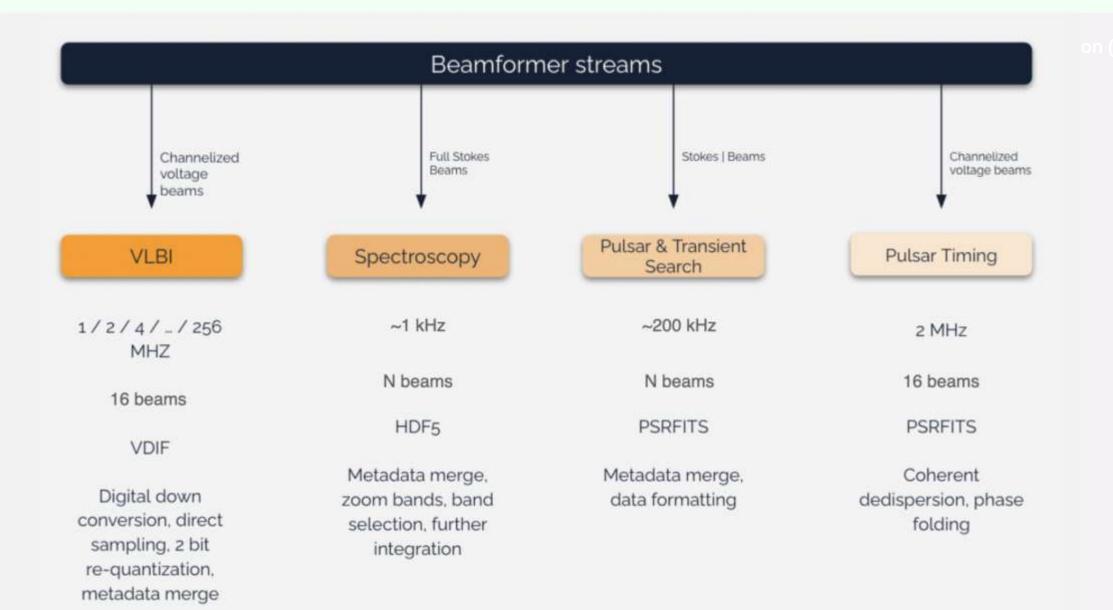
Current status: Factory acceptance tests for the first prototype



Backend / Correlator



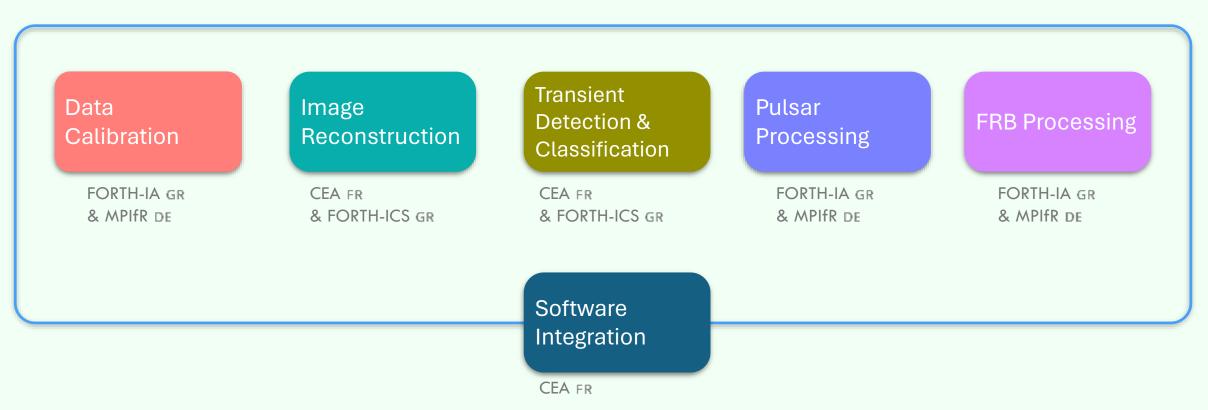
Backend / Correlator



Science Pipelines

S COSMOSTAT 🤷 ARG S

Current Development effort

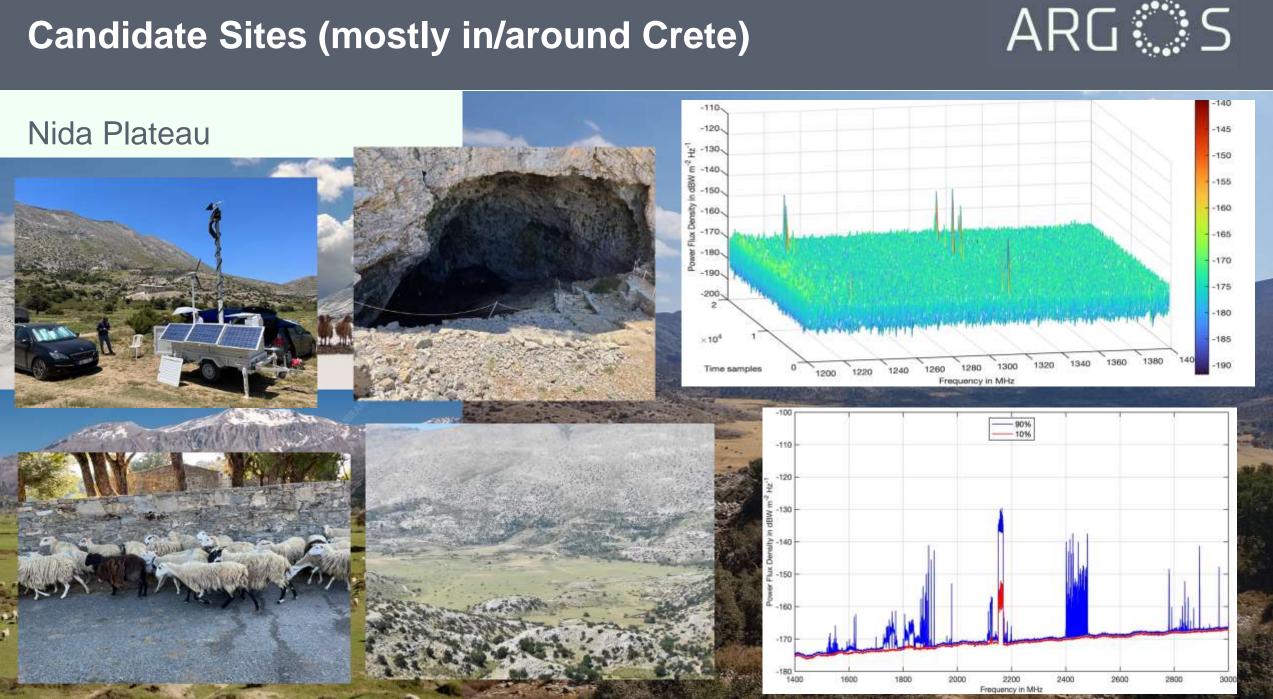


Candidate Sites (mostly in/around Crete)





Candidate Sites (mostly in/around Crete)



Join us!





Several ways; start here http://argos.ia.forth.gr/ and follow the links

- Sign up for the newsletter
- Participate in the Science Working Group (john@ia.forth.gr)
- workshops (TBD, Q2 2025 in Crete)
- White Paper (2024)
- Apply for jobs





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Max-Planck-Institut für Radioastronomie





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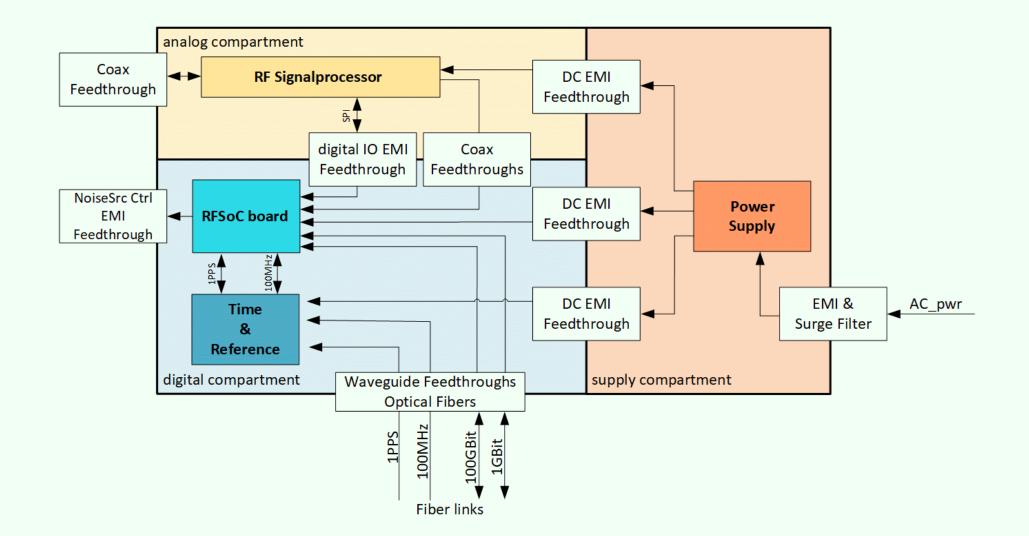
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Frontend and Digital Feed





WP7 Signal Processing

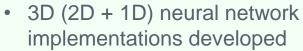
Wavelets time axis Normalized & centered reconstruction $w_{j_1,j_2}[k_x,k_y,k_z]$ DATA $= D * \overline{\psi}_{j_1}^{(xy)} * \overline{\psi}_{j_2}^{(z)}(x, y, z)$ 2D1DCLEAN $N_p = N_{2D} Nx Ny Nz$ 2D Wavelet 2D Wavelet 2D Wordet L-band (1.4 GHz) Pilia et al. 2015 WTID WTID WTID 0.3 0.4 0.5 0.6 0.7 Pulsar fractional period Neural Networks CLEAN 0.200 Encoder(y₀) SFTNet(_, h₀, σ_c) 2D-1D Net 0.175 **Deflation Net** 0.150 9.0.125 Sneoder(gan) SFTNet(., h₃₁, σ_e) 0.100 0.075 | AVG 0.050 0.025 Encoder(ii) -SFTNet(_h 0.000

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- 2D and 3D image reconstruction software developed
- Outperforms community standard techniques such as CLEAN
- Needs to be benchmarked on ARGOS simulations



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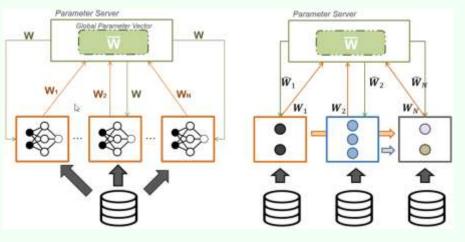


- Outperforms community standard techniques such as CLEAN
- Paper published by Chiche et al. (2023)

WP8: Archive, Alerts, UI



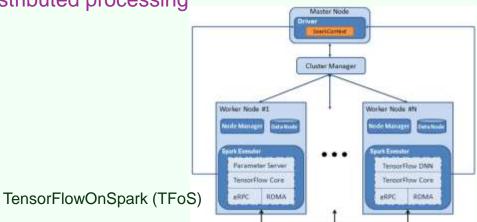
GPU-based architectures for accelerated DL execution & optimal management of cluster resources



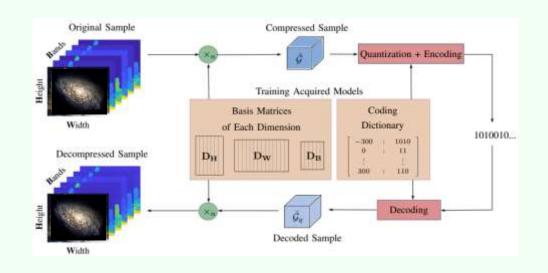
Data parallelization

Model parallelization

Services @ Apache Spark for large-scale data storage & distributed processing



Tensor decomposition learning for compression & recovery of high-dimensional signals



WP9: Dissemination, exploitation, communication

ARG

Dissemination in numbers

- 12 peer-reviewed papers in 2023
- 15 invited talks and colloquia
- 3 press releases
- 2 MSc theses
- 3 (+1) PhD students
- 4 BSc student theses (in prep.)



DIALTHMA 25 AUGUSTUS 2023 (1715

ΑΡΓΟΣ: Εξασφαλίστηκαν τα πρώτα κονδύλια για νέο ραδιοτηλεσκόπιο στον Ψηλορείτη

Στην πιλοτική φάση θα εγκατασταθούν στο Αστεροσκοπείο του Σκίνακα 16 παραβολικές κεραίες.

nature

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Monster gravitational waves spotted for first time

....



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Scientists pick up shock waves from colliding galaxies





Έξι διαφορετικά τηλεσκόπια στην Ευρώπη επιβεβαιώνουν: Τα βαρυτικά κύματα υπάρχουν

Νέο εποτημονικά δεδομένο στον δρόμο τρος την εξιενίσας της δημιωργίας του Σύμταντος, με τη συμμετατή του Ινατιτούτου Αστροφισικής του ITE

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Ραδιοτηλεσκόπιο "ΑΡΓΟΣ": Το ΙΤΕ ανοίγει ένα νέο παράθυρο στο Σύμπαν

Ευρωπαϊκής σύμπραξης ηγείται το Ινστιτούτο Αστροφυσικής του ΙΤΕ - Σημαντική ευρωπαϊκή χρηματοδότηση 3 εκ. ευρώ