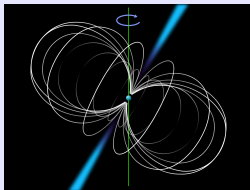


The structure of strange stars with a new quark mass scaling

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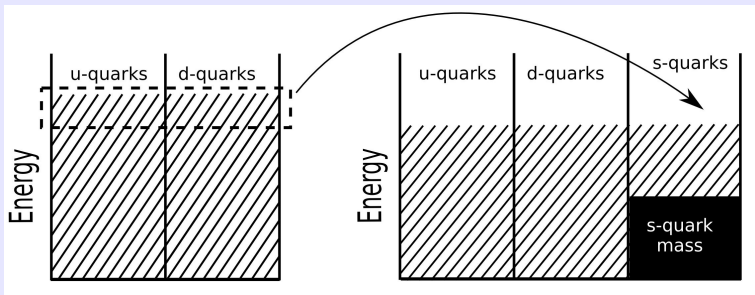
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- 3 Properties of SQM and strange stars
- 4 Summary and Outlook

Strange Quark Matter (SQM)



Robert Jensen, Searches for Strange Quark Matter, March 2006

- Bodmer first suggested a low energy nuclear state called “**collapsed nuclei**” . (Bodmer1971_PRD4-1601)
- Witten reported on **the stability of strange quark matter (SQM)** consisting of approximately equal numbers of u , d and s quarks, suggesting that SQM could indeed be stable even at zero external pressure. (Witten1984_PRD30-272)

Phenomenological models

In principle, the properties of SQM can be studied based on **quantum chromodynamics (QCD)**. However, due to the known **difficulties** in the nonperturbative region, **phenomenological models** are essential for the study of SQM, e.g.

MIT bag model: The vacuum has a **constant energy density**, that is, the bag constant B provides a negative pressure to **confine** quarks.

Equiparticle model: The **strong interaction** is considered by adopting **equivalent quark masses** while the free energy density and particle number densities have the same form as a free particle system.

Other models: Nambu and Jona-Lasinio (NJL) model; Perturbation model; Quark-cluster model; Quasiparticle model; Global color symmetry model (GCM); Field correlator method;

...

Equiparticle model

Quasiparticle approach: The medium created in ultrarelativistic nucleus-nucleus collisions **interacts more strongly than hadron or string matter**. (Peshier_Cassing2005_PRL94-172301)

① The confinement is automatically achieved without an additional bag constant;

② The mass scaling is related to the in-medium chiral condensate (Peng_Chiang_Yang_Li_Liu1999_PRC61-015201):

$$m_i = \frac{E_i}{\sum_j ((\bar{q}_i q_i) - \langle \bar{q}_i q_i \rangle_0)}$$
(1)

③ The results from more fundamental approaches (Lattice QCD and Perturbative QCD) can be incorporated.

Note: Self-consistent thermodynamic treatment is required.

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$$m_{\Gamma} = \frac{E_{\Gamma}}{\sum_i (\langle \bar{q}_i q_i \rangle - \langle \bar{q}_i q_i \rangle_0)}; \quad (1)$$

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Thermodynamics of equiparticle model

The equivalent mass for quark flavor i is $m_i = m_i(n_u, n_d, n_s, T)$. For given T , V and n_i , the free energy density is then $F = F(T, V, \{n_i\}, \{m_i\})$.

Note that F and n_i have the same form as a **free particle system**.

Standard thermodynamics:

$$dF = -SdT + \left(-P - F + \sum_i \mu_i n_i\right) \frac{dV}{V} + \sum_i \mu_i dn_i. \quad (2)$$

Equiparticle model:

$$dF = \left[\frac{\partial \Omega_0}{\partial T} + \sum_i \frac{\partial \Omega_0}{\partial m_i} \frac{\partial m_i}{\partial T} \right] dT + \frac{\partial \Omega_0}{\partial V} dV + \sum_i \left[\mu_i^* + \sum_j \frac{\partial \Omega_0}{\partial m_j} \frac{\partial m_j}{\partial n_i} \right] dn_i. \quad (3)$$

Entropy density: $S = -\frac{\partial \Omega_0}{\partial T} - \sum_i \frac{\partial \Omega_0}{\partial m_i} \frac{\partial m_i}{\partial T}$;

Pressure: $P = -F + \sum_i \mu_i n_i - V \frac{\partial \Omega_0}{\partial V}$;

Chemical potential: $\mu_i = \mu_i^* + \sum_j \frac{\partial \Omega_0}{\partial m_j} \frac{\partial m_j}{\partial n_i}$.



C.-J. Xia, G.-X. Peng, S.-W. Chen, Z.-Y. Lu & J.-F. Xu
Phys. Rev. D, 2014, 89,
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History of quark mass scaling

Equivalent mass for quark flavor i : $m_i = m_{i0} + m_1$.

- Originally it was suggested to be an **inversely linear scaling**: $m_1 = \frac{B}{3n}$; (Fowler_Raha_Weiner1981_ZPC9-271)
- Based on this scaling, in 2002 Zhang and Su included finite temperature in the scaling: $m_1 = \frac{B}{3n} \left[1 - \left(\frac{T}{T_c} \right)^2 \right]$; (Zhang_Su2002_PRC65-035202)
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Adopting the scaling method, Peng et al. extended the scaling to include

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- 6 The **quark matter symmetry energy** was considered by Chu and Chen: $m_I = \frac{D}{n^{1/3}} - \tau_1 \delta D_I n^\alpha e^{-\beta \alpha}$; (Chu_Chen2014_ApJ780-135)

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A new quark mass scaling

Adopting the similar method of Peng2005_NPA747-75, namely expanding the equivalent mass to a **Laurant series** of the holistic Fermi momentum ν , and take the leading term in both directions:

$$m_I = \frac{a_{-1}}{\nu} + a_1 \nu, \quad (4)$$

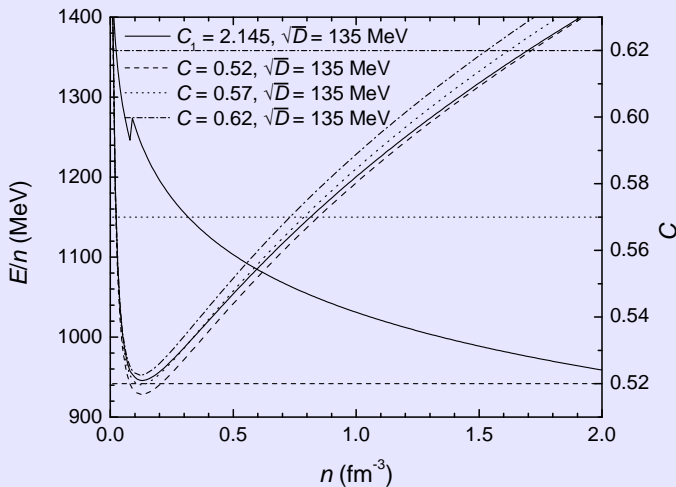
here the first term corresponds to the **linear confinement**, while the second term is responsible for the leading-order **perturbative interactions**. The mass scaling is then given by

$$m_i = m_{i0} + \frac{D}{n^{1/3}} + Cn^{1/3}, \quad (5)$$

where $C = C_1 a_1 \approx C_1 \sqrt{\frac{2}{3}} \alpha$. Since the strong coupling runs **logarithmically**, the running rate is thus much slower and the parameter C can be taken as **constant**. According to the analytic coupling constant, the maximum value of α is $1/\beta_0$, then we have

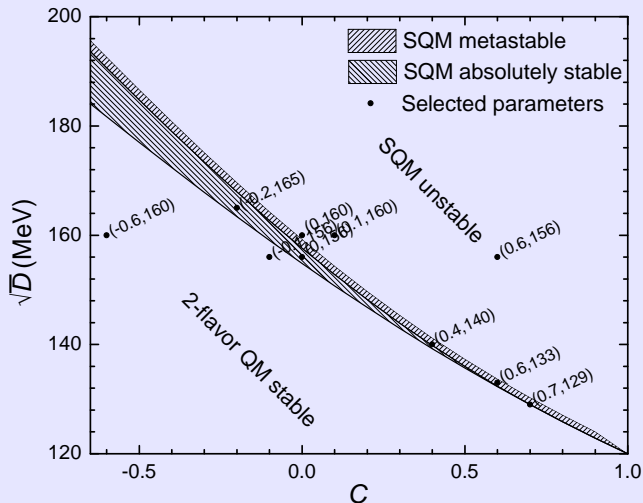
$$C < \left(\frac{3\pi^2}{N_f} \right)^{1/3} \sqrt{\frac{2}{3\beta_0}} \approx 1.1676. \quad (6)$$

Variation range of C



- 1 The fluctuation of the energy per baryon is **small with running coupling;**
- 2 The one-gluon-exchange interaction reduces C at lower densities.

The range of the parameters D and C



- 1 The current quark masses are taken as $m_{u0} = 5$ MeV, $m_{d0} = 10$ MeV and $m_{s0} = 100$ MeV;
- 2 The solid dots are the combinations of parameters we are about to use.

Calculate the properties of SQM

The weak equilibrium conditions:

$$\mu_u + \mu_e = \mu_d = \mu_s. \quad (7)$$

The charge neutrality condition:

$$\frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s - n_e = 0. \quad (8)$$

The baryon number conservation:

$$n = \frac{1}{3}(n_u + n_d + n_s). \quad (9)$$

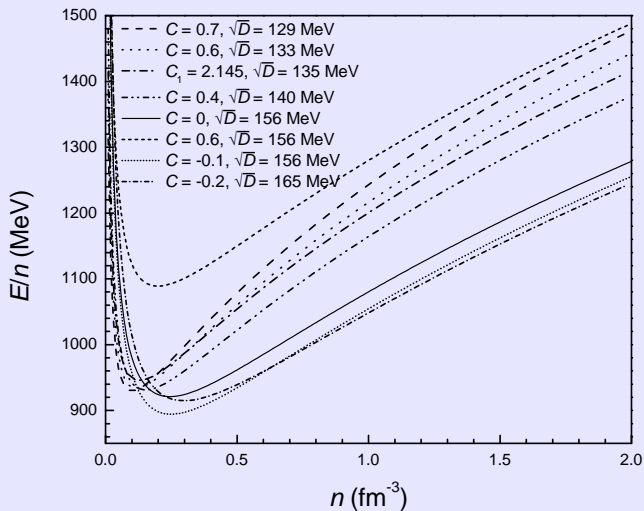
Pressure and chemical potential

Based on the newly obtained mass scaling (5), taking the temperature $T = 0$ and volume $V \rightarrow \infty$, the pressure and chemical potential can be obtained by

Pressure:
$$P = -\Omega_0 + n \frac{dm_I}{dn} \frac{\partial \Omega_0}{\partial m_I};$$

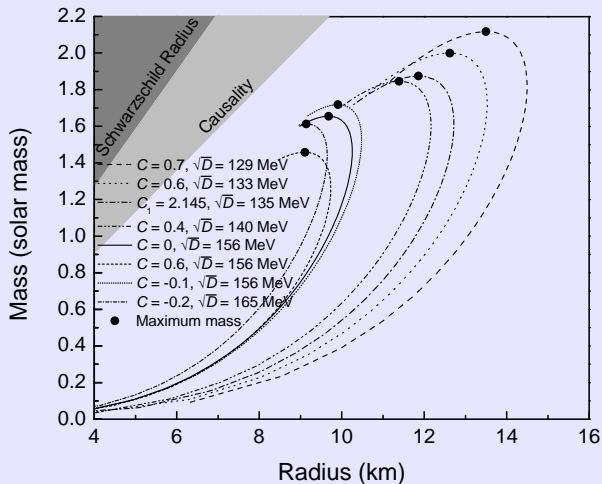
Chemical potential:
$$\mu_i = \mu_i^* + \frac{1}{3} \frac{dm_I}{dn} \frac{\partial \Omega_0}{\partial m_I}.$$

Energy per baryon



- 1 The EOS becomes **stiffer** for parameters within the “SQM stable” area;
- 2 Increasing C or decreasing D also makes the EOS stiffer;
- 3 **Zero pressure** point can be smaller than nuclear saturation density, and quark hadron **phase transition** will occur.

Mass-radius relation of strange stars



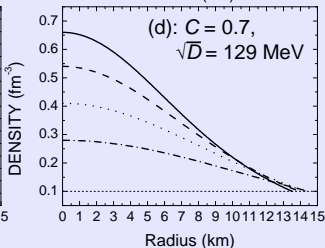
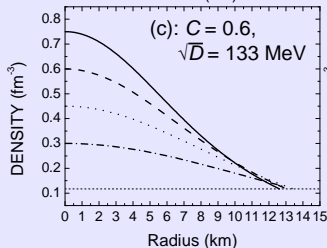
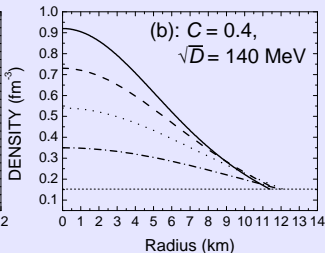
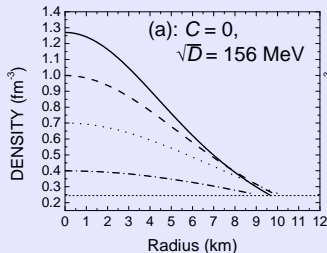
To obtain the structure of strange stars, we have used the Tolman-Oppenheimer-Volkov (TOV) Equation:

$$\frac{dP}{dr} = -(P+E) \frac{G}{r^2} \frac{4\pi r^3 P + M}{1 - \frac{2GM}{r}}$$

with

$$M(r) = \int_0^r 4\pi E r^2 dr.$$

Density profiles



- 1 The uppermost curve corresponds to the largest acceptable central density;
- 2 The horizontal line corresponds to the surface density of the star;
- 3 The surface density gets even lower than nuclear saturation density, and the quark hadron phase transition should be considered.

Summary and Outlook

Summary

- 1 A **new quark mass scaling** with linear confinement and leading-order perturbative interactions is obtained by expanding the equivalent mass to a Laurant series and taking the leading terms in both directions;
- 2 With the new quark mass scaling and self-consistent thermodynamic treatment, we have studied **the equation of state (EOS) of SQM** with various combination of parameters;
- 3 Based on the EOS of SQM and TOV equation, we studied **the structure of strange stars**, where the masses and radii of PSR J1614-2230 and PSR J0348+0432 can be reproduced.

Outlook

- 1 Quark hadron phase transition and hybrid stars;
- 2 The effect of electric field on the structure of compact stars;
- 3 The effect of magnetic field on the structure of compact stars;
- 4 Compact stars at finite temperature with nonzero neutrino chemical potential;
- 5 ...

Thank You!!!